

### Radial Potential

energy and angular momentum

$$e \equiv \frac{E}{m} = \frac{1}{2} [\dot{r}^2 + r^2 \dot{q}^2] - \frac{GM}{r}$$

$$L \equiv \frac{J}{m} = r^2 \dot{q}$$

effective potential  $e = \frac{\dot{r}^2}{2} + \Phi(r)$

$$\Phi(r) = \frac{L^2}{2r^2} - \frac{GM}{r} = e_0 \left( \frac{\ell^2}{r^2} - 2 \frac{\ell}{r} \right)$$

circular orbit:

$$\frac{\partial \Phi}{\partial r} = \frac{GM}{r^2} - \frac{L^2}{r^3} = 0 \rightarrow r = \frac{L^2}{GM} \equiv \ell$$

$$\Phi(\ell) = -e_0 \quad e_0 \equiv \frac{GM}{2\ell} = \frac{1}{2} \left( \frac{GM}{L} \right)^2 = \frac{1}{2} \left( \frac{L}{\ell} \right)^2$$

Note:  $\ell, e_0$  depend on  $M$  and  $J$  but not  $E$ .

AS 4024 Binary Stars and Accretion Disks

### Types of Orbits

Fix  $L, E = \min$  circular  
 $E < 0$  bound (ellipse)  
 $E > 0$  unbound (hyperbola)

Fix  $E < 0, L = \max$  circular  
 $L = \min$  radial

AS 4024 Binary Stars and Accretion Disks

### Turning Points

energy and angular momentum:

$$e = \frac{\dot{r}^2}{2} + \frac{L^2}{2r^2} - \frac{GM}{r}$$

$$= \frac{\dot{r}^2}{2} + e_0 \left[ (u-1)^2 - 1 \right]$$

$$u \equiv \frac{\ell}{r}$$

turning points:

$$\dot{r}^2 = 0 \rightarrow (u-1)^2 = 1 + \frac{e}{e_0} \equiv e^2$$

$$u^{\pm} = \frac{\ell}{r^{\pm}} = 1 \pm \left( 1 + \frac{e}{e_0} \right)^{1/2}$$

AS 4024 Binary Stars and Accretion Disks

### Orbit Shape

$$e = \frac{\dot{r}^2}{2} + e_0(x^2 - 1)$$

$$r^2 = 2e_0 \left( \frac{e}{e_0} + 1 - x^2 \right) = \left( \frac{L}{\ell} \right)^2 (e^2 - x^2)$$

$$\frac{dx}{dq} = \frac{\dot{x}}{\dot{q}} = \left( -\frac{\ell \dot{r}}{r^2} \right) / \left( \frac{L}{r^2} \right) = -\frac{\ell}{L} \dot{r}$$

$$\left( \frac{dx}{dq} \right)^2 + x^2 = e^2$$

simple harmonic oscillator!

$$x = e \cos q \rightarrow r = \frac{\ell}{1 + e \cos q}$$

AS 4024 Binary Stars and Accretion Disks

### Conic Sections

$$r = \frac{\ell}{1 + e \cos q}$$

$e$  = eccentricity  
 $> 1$  hyperbola  
 $= 1$  parabola  
 $< 1$  ellipse  
 $= 0$  circle

$\ell$  = semi-latus rectum  $= a(1 - e^2)$   
 $a$  = semi-major axis  
 $b$  = semi-minor axis  
 $= a(1 - e^2)^{1/2}$

$q = 0^\circ \quad r = \frac{\ell}{1+e} = a(1-e)$  periastron  
 $q = 90^\circ \quad r = \frac{\ell}{1+e} = a(1-e^2)$   
 $q = 180^\circ \quad r = \frac{\ell}{1-e} = a(1+e)$  apastron

AS 4024 Binary Stars and Accretion Disks

### Orbital Speed

$$r = \frac{\ell}{1 + e \cos q} \quad \ell \equiv \frac{L^2}{GM} = a(1 - e^2)$$

$$e = -\frac{GM}{2\ell} (1 - e^2) = -\frac{GM}{2a}$$

Energy:

$$v^2 = 2e + \frac{2GM}{r} = GM \left[ \frac{2}{r} - \frac{1}{a} \right]$$

AS 4024 Binary Stars and Accretion Disks

### Motion in Time

$$r = \frac{\ell}{1 + e \cos q} \quad \ell \equiv \frac{L^2}{GM}$$

$$\dot{r} = \frac{L}{r^2} = \frac{L}{\ell^2} (1 + e \cos q)^2$$

$$\frac{dq}{(1 + e \cos q)^2} = \frac{L}{\ell^2} dt$$

$$\int \frac{dq}{(1 + e \cos q)^2} = \frac{L}{\ell^2} (t - T)$$

No analytic solution for  $q(t)$

AS 4024

Binary Stars and Accretion Disks

### Eccentric Anomaly

$$b = a\sqrt{1 - e^2}$$

$q$  = true anomaly

$E$  = eccentric anomaly

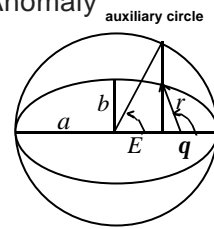
$$x = a \cos E = a e + r \cos q$$

$$y = b \sin E = r \sin q$$

$$r^2 = a^2 (\cos^2 E - e)^2 + a^2 (1 - e^2) \sin^2 E$$

$$r = a (1 - e \cos E) \rightarrow dr = a e \sin E dE$$

.....Kepler's equation giving  $E(t)$



AS 4024

Binary Stars and Accretion Disks

### Motion in Time

$h$  = mean anomaly

$f$  = orbital phase

$P$  = orbital period

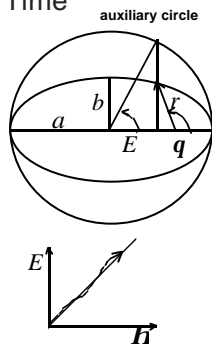
$T$  = time of periastron passage

Kepler's equation :

$$E - e \sin E = h = 2\pi f = \frac{2\pi}{P} (t - T)$$

iterate to find  $E(t)$

$$\tan\left(\frac{q}{2}\right) = \left(\frac{1+e}{1-e}\right)^{1/2} \tan\left(\frac{E}{2}\right)$$



AS 4024

Binary Stars and Accretion Disks