

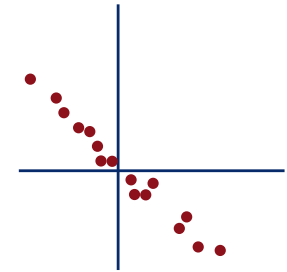
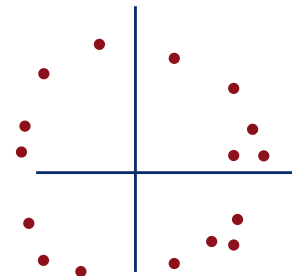
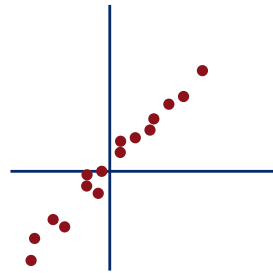
# Review: Correlation vs Independence

•  $R > 0$

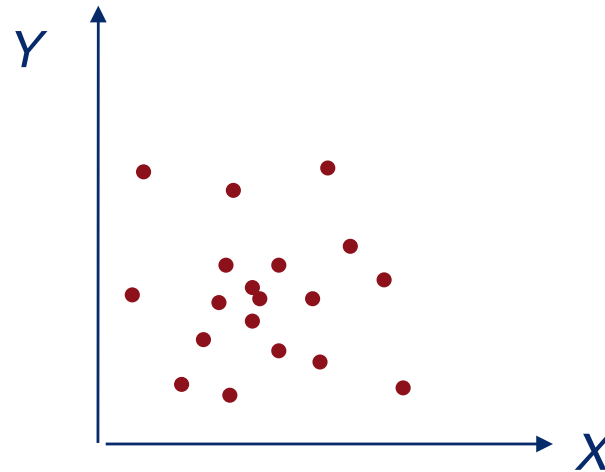
•  $R = 0$

•  $R < 0$

Dependent :

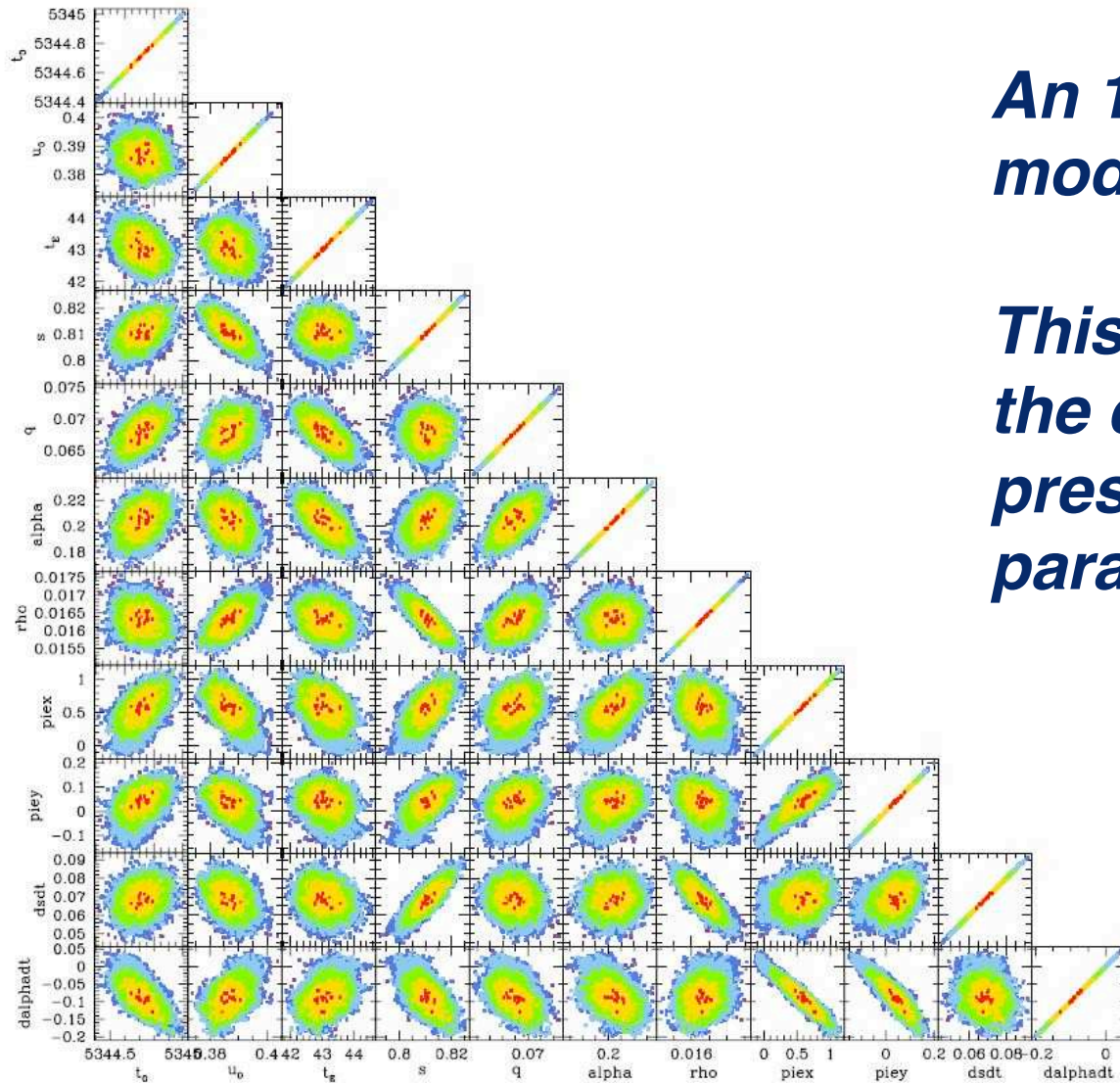


Independent :



$$f(X, Y) = f(X) f(Y)$$

# Example: Correlation Matrix



***An 11-parameter model fitted to data.***

***This matrix shows the correlations present among the 11 parameters.***

# Review: Algebra of Random Variables

$$\langle a \rangle = a$$

$$\text{Var}[a] = 0$$

$$\langle aX \rangle = a\langle X \rangle$$

$$\text{Var}[aX] = a^2 \text{Var}[X]$$

$$\langle X + Y \rangle = \langle X \rangle + \langle Y \rangle$$

$$\text{Var}[X + Y] = \text{Var}[X] + \text{Var}[Y] + 2\text{Cov}[X, Y]$$

Co - variance :

$$\text{Cov}[X, Y] \equiv \langle (X - \langle X \rangle)(Y - \langle Y \rangle) \rangle \quad \text{Var}[X] \equiv \text{Cov}[X, X]$$

Linear transformations :

$$\left\langle \sum_i a_i X_i \right\rangle = \sum_i a_i \langle X_i \rangle$$

$$\text{Var}\left[ \sum_i a_i X_i \right] = \sum_i \sum_j a_i a_j \sigma_i \sigma_j R_{ij}$$

Correlation Matrix :

$$R_{ij} \equiv \frac{\text{Cov}(X_i, X_j)}{\sigma_i \sigma_j}$$

$$\sigma_i \equiv \sigma(X_i)$$

# Practice the “fuzzy” algebra of random variables

$$6 (1 \pm 1) =$$

$$(1 \pm 1) + (2 \pm 2) =$$

$$(1 \pm 2) - (2 \pm 2) =$$

*Practice until this becomes automatic ...*

# Functions of Random Variables

Often what we can measure is not what we are most interested in!

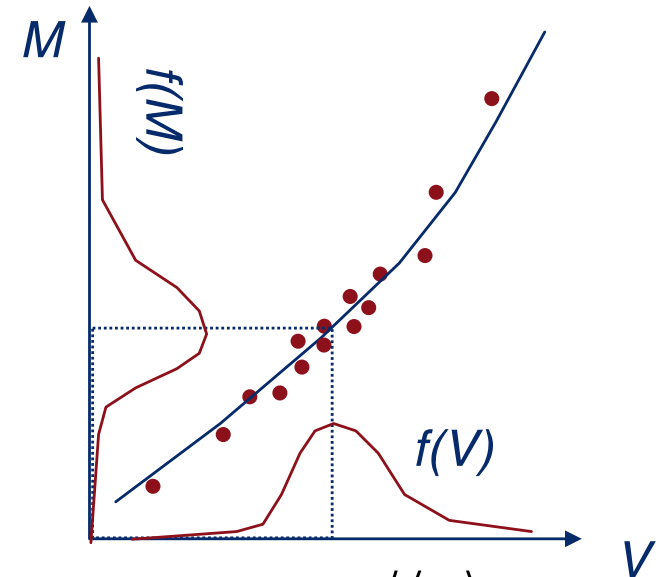
Example: mass of binary-star system:

$$M = \frac{V^2 a}{G} = \frac{V^3 P}{2\pi G}$$

We want  $M$ , but can only measure  $V$  and  $P$ .

$P$  = accurate, but  $V$  usually less certain.

What is the uncertainty in  $M$ ?



For power-laws:  $\ln M = 3 \ln V + \ln P + \text{const.}$   $\sigma(\ln x) \approx \sigma(x)/\langle x \rangle$

$$\left( \frac{\sigma_M}{\langle M \rangle} \right)^2 \approx \left( 3 \frac{\sigma_V}{\langle V \rangle} \right)^2 + \left( \frac{\sigma_P}{\langle P \rangle} \right)^2$$

(valid for *small* and *independent* errors in  $V$  and  $P$ ).

**How do error bars propagate through non-linear functions?**

# Functions of a Random Variable

$$Y = y(X) \quad \frac{dY}{dX} = y'(X)$$

conserve probability:

$$d(\text{Prob}) = f(Y) |dY| = f(X) |dX|$$

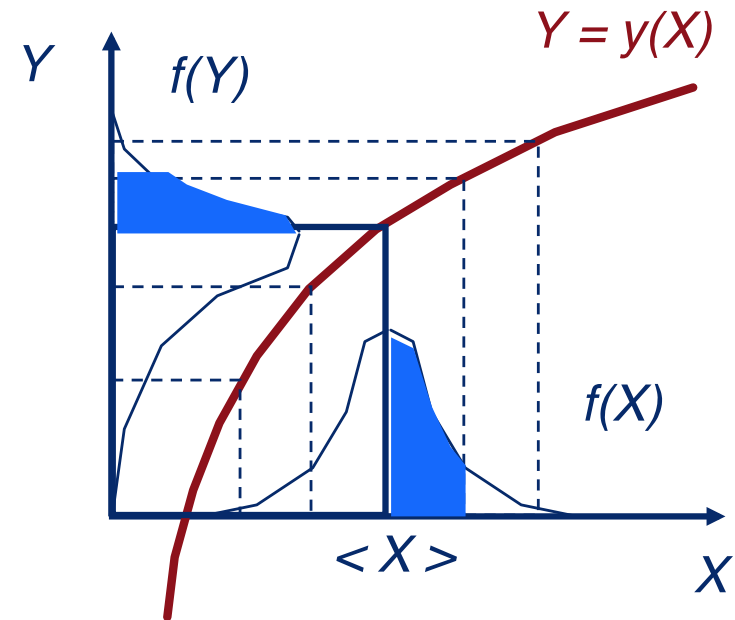
$$f(Y) = f(X) \left| \frac{dX}{dY} \right| = \frac{f(X)}{|y'(X)|}$$

mean value (biased)

$$\langle Y \rangle = y(\langle X \rangle) + \frac{1}{2} y''(\langle X \rangle) \sigma_X^2 + \dots$$

standard deviation (stretched)

$$\sigma_Y = \sigma_X \left| \frac{dy}{dx} \right|_{X=\langle X \rangle} + \dots$$



Negative curvature:

Long tail for  $Y < y(\langle X \rangle)$

Bias:  $\langle Y \rangle < y(\langle X \rangle)$ .

Median is not biased:  
 $\text{Med}(Y) = y(\text{Med}(X))$

# Examples of Non-linear Transformations

Spectral Energy Distributions: per unit **wavelength** ( $\text{erg cm}^{-2} \text{A}^{-1}$ ),  
or per unit **frequency** ( $\text{erg cm}^{-2} \text{s}^{-1} \text{Hz}^{-1}$ )

$$f_\nu(\lambda) |d\nu| = f_\lambda(\lambda) |d\lambda|$$

$$\nu = \frac{c}{\lambda} \quad d\nu = -\frac{c}{\lambda^2} d\lambda \quad \Rightarrow \quad f_\nu(\lambda) = \left| \frac{d\lambda}{d\nu} \right| f_\lambda(\lambda) = \frac{\lambda^2}{c} f_\lambda(\lambda)$$

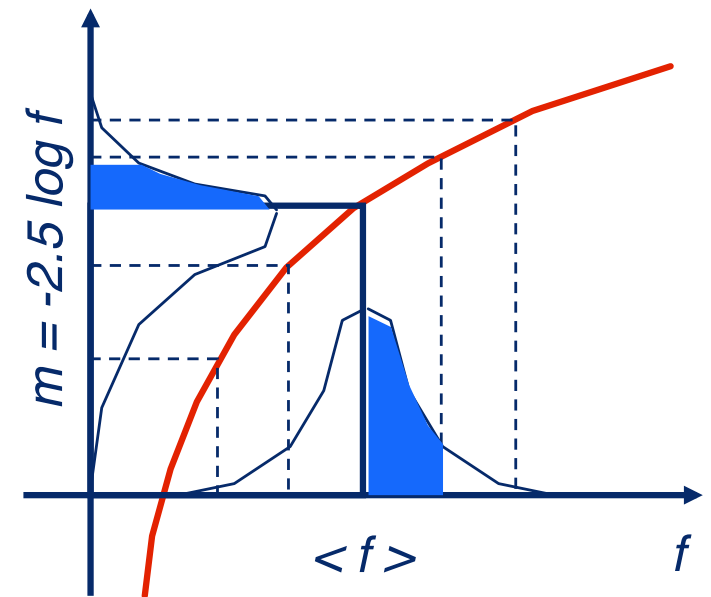
## Converting a **flux** to a **magnitude**:

- Measure Flux: Gaussian distribution:  $f \sim G(\langle f \rangle, \sigma_f^2)$
- Nonlinear transformation induces a bias:

$$m = m_0 - 2.5 \log f$$

$$\langle m \rangle = m_0 - 2.5 \log \langle f \rangle + a \sigma_m^2$$

- PROBLEM: evaluate  $a, \sigma_m$  in terms of  $\langle f \rangle, \sigma_f$ .



# Nonlinear Transformations: A Bias from Curvature + Noise

Taylor expand  $Y = y(X)$  around  $X = \langle X \rangle$ :

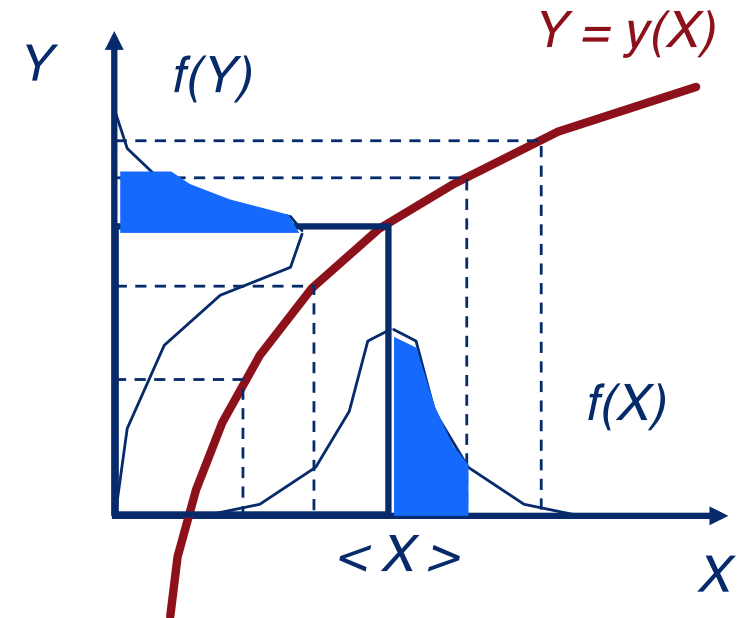
$$y(X) = y(\langle X \rangle) + y'(\langle X \rangle) \varepsilon + \frac{1}{2} y''(\langle X \rangle) \varepsilon^2 + \dots$$

where  $\varepsilon \equiv X - \langle X \rangle$ ,  $\langle \varepsilon \rangle = 0$ ,  $\langle \varepsilon^2 \rangle = \sigma_X^2$ .

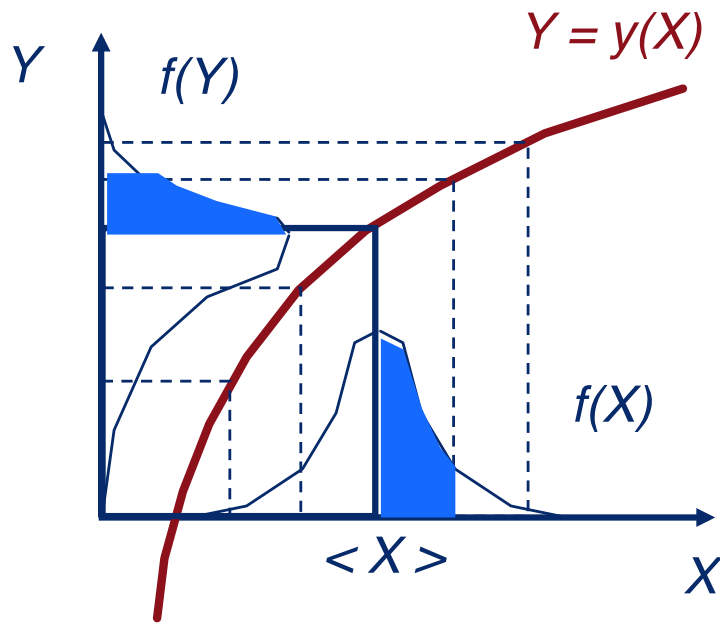
Hence

$$\begin{aligned} \langle y(X) \rangle &= \left\langle y(\langle X \rangle) + y'(\langle X \rangle) \varepsilon + \frac{1}{2} y''(\langle X \rangle) \varepsilon^2 + \dots \right\rangle \\ &= y(\langle X \rangle) + y'(\langle X \rangle) \langle \varepsilon \rangle + \frac{1}{2} y''(\langle X \rangle) \langle \varepsilon^2 \rangle + \dots \\ &= y(\langle X \rangle) + 0 + \frac{1}{2} y''(\langle X \rangle) \sigma_X^2 + \dots \end{aligned}$$

**This is the bias.**



# Variance of a Transformed Variable



**Tangent-curve approximation :**

$\sigma( y(x) ) = \sigma(x)$  **stretched** by a factor  $|dy/dx|$ .

$$\begin{aligned} \sigma^2(Y) &\equiv \langle (Y - \langle Y \rangle)^2 \rangle = \left\langle \left[ y(\langle X \rangle) + y'(\langle X \rangle) \varepsilon + \frac{1}{2} y''(\langle X \rangle) \varepsilon^2 + \dots \right. \right. \\ &\quad \left. \left. - y(\langle X \rangle) - 0 - \frac{1}{2} y''(\langle X \rangle) \sigma^2(X) - \dots \right]^2 \right\rangle \\ &= \left\langle \left[ y'(\langle X \rangle) \varepsilon + O(\varepsilon^2) \right]^2 \right\rangle = [y'(\langle X \rangle)]^2 \sigma_X^2 \end{aligned}$$

Can extend to higher-order terms (skew, kurtosis) if needed.

# Magnitude Bias

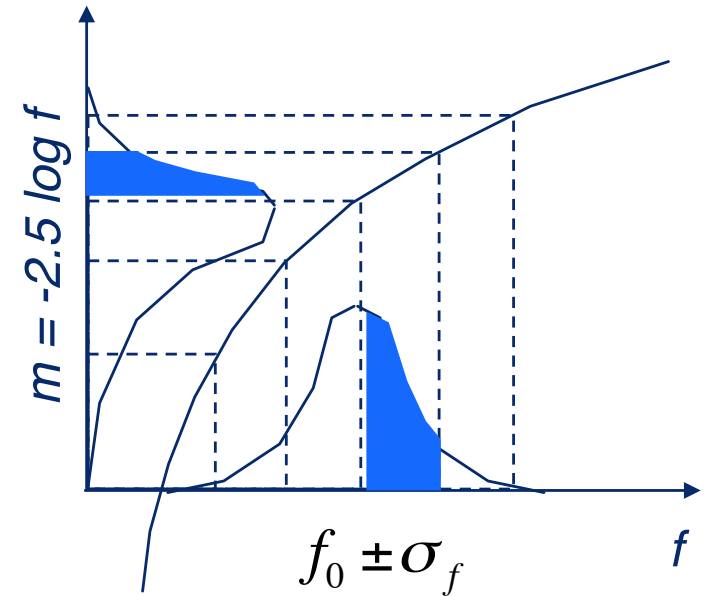
Observe flux:  $f \sim G(f_0, \sigma_f^2) = (f_0 \pm \sigma_f)$

Transform to magnitude:  $m(f) \equiv m_0 - 2.5 \log f$

Derivatives: ( $\log f = \log e \ln f$ )  $\frac{dm}{df} = -\frac{2.5 \log e}{f}$ ,  $\frac{d^2m}{df^2} = \frac{2.5 \log e}{f^2}$ .

$$\sigma_m \approx \left| \frac{dm}{df} \right|_{\langle f \rangle} \sigma_f = \frac{2.5 \log e}{\langle f \rangle} \sigma_f \approx 1.08 \frac{\sigma_f}{\langle f \rangle}.$$

$$\begin{aligned} \langle m \rangle &= m(\langle f \rangle) + \frac{1}{2} \frac{d^2m}{df^2} \Big|_{\langle f \rangle} \sigma_f^2 \\ &= m_0 - 2.5 \log \langle f \rangle + \frac{2.5 \log e}{2 \langle f \rangle^2} \sigma_f^2 \\ &= m_0 - 2.5 \log \langle f \rangle + \frac{\sigma_m^2}{5 \log e} \end{aligned}$$



Note the bias toward faint magnitudes.

# Magnitude Bias

$$m(f) \equiv m_0 - 2.5 \log f$$

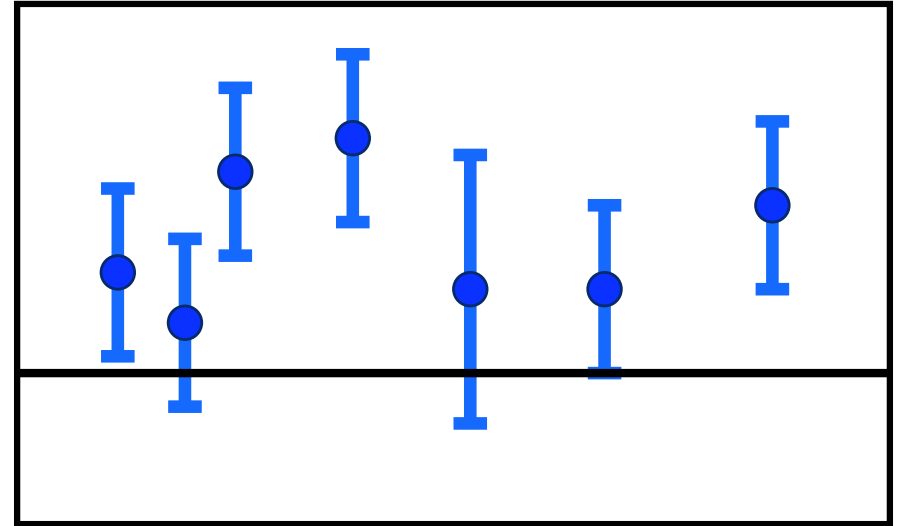
$$\sigma_m = (2.5 \log e) \frac{\sigma_f}{\langle f \rangle} \approx 1.08 \frac{\sigma_f}{\langle f \rangle}.$$

$$\langle m \rangle = m(\langle f \rangle) + \text{bias}(\langle m \rangle)$$

$$\text{bias}(\langle m \rangle) = \left( \frac{\sigma_m}{\sqrt{5 \log e}} \right)^2 \approx 0.01 \left( \frac{\sigma_m}{0.15} \right)^2$$

**15% errors -> 1% bias**

**50% errors -> 10% bias**



**Given noisy fluxes, should you average fluxes and then compute the magnitude?**

**Or, should you convert fluxes to magnitudes and average the magnitudes ?**

$$m(f) = m_0 - 2.5 \log \langle f \rangle$$

$$\text{or } m(f) = \langle m_0 - 2.5 \log f \rangle$$

# Cartesian $\rightarrow$ Polar coordinates e.g. Amplitude and Phase

Independent measurements of  $C$  and  $S$

( e.g. cos and sin amplitudes of an oscillation):

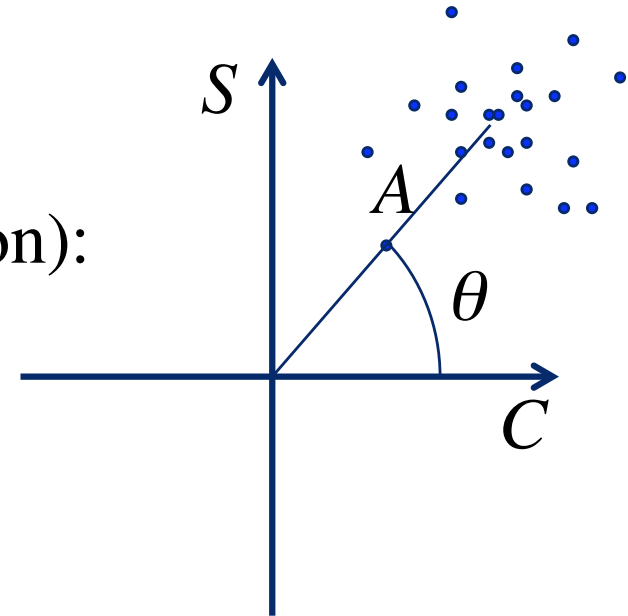
$$S = A \sin \theta \sim (S_0 \pm \sigma_S)$$

$$C = A \cos \theta \sim (C_0 \pm \sigma_C)$$

Transform to amplitude and phase:

$$A = ? \pm ?$$

$$\theta = ? \pm ?$$



# Transforming Random Numbers

Uniform  $\rightarrow$  Lorentzian

$$u \sim U(0,1) \rightarrow x \sim L(\mu, \sigma)$$

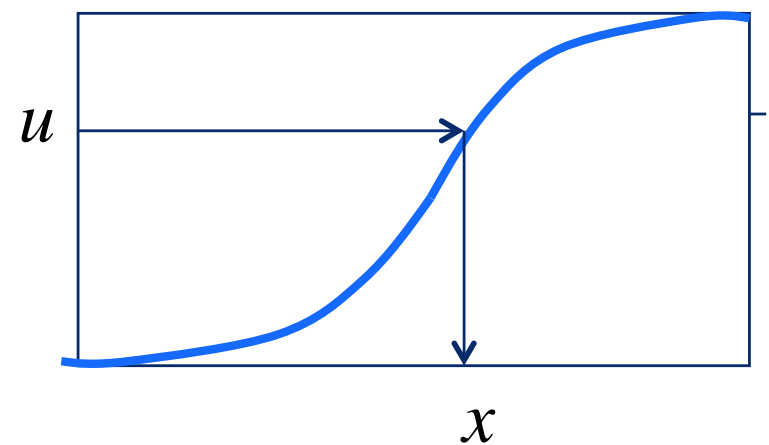
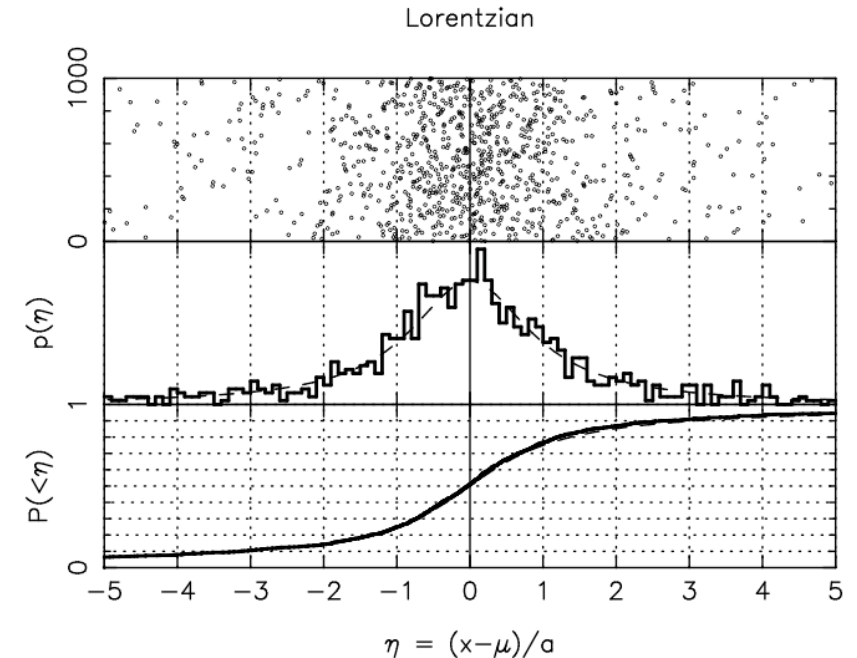
$$u = F(x) = \frac{1}{\pi} \arctan\left[\frac{x - \mu}{\sigma}\right] + \frac{1}{2}$$

$$x = F^{-1}(u) = \mu + \sigma \tan\left[\pi\left(u - \frac{1}{2}\right)\right]$$

Practice :

Uniform  $\rightarrow$  Exponential

Uniform  $\rightarrow$  Power - law



# Box-Muller Transform

For Gaussians, cumulative probability  $F(x)$  has no analytic expression.  
Harder to generate Gaussian random numbers  $x = F^{-1}(u)$   
from Uniform random numbers  $u$ .

Two independent uniform random numbers :

$$x \sim U(-1, +1) \quad y \sim U(-1, +1)$$

Keep if  $r^2 = x^2 + y^2 < 1$

Two independent gaussian random numbers :

$$G_1 = \frac{2x}{r} (-\ln r)^{1/2} \quad G_2 = \frac{2y}{r} (-\ln r)^{1/2}$$

$$r = 0 \quad \rightarrow \quad G = \infty$$

$$r = 1 \quad \rightarrow \quad G = 0$$

$G_1$  and  $G_2$  have mean 0 and variance 1:

$$G_1 \sim G(0,1) \quad G_2 \sim G(0,1)$$

