

"The Origins of the Molecular Building Blocks of Life"

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Astrobiologists ask...

Whether there is life elsewhere in
our universe...

(by concentrating (right now) on
finding suitable habitable zones,
exoplanets, prebiotic molecules &
using solar system as lab)



Star Formation

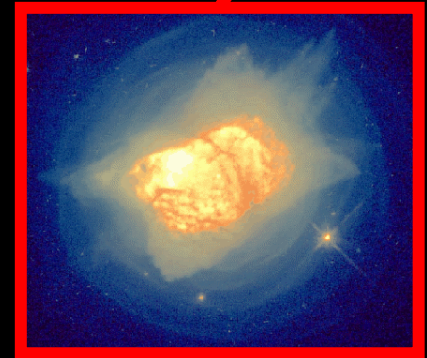
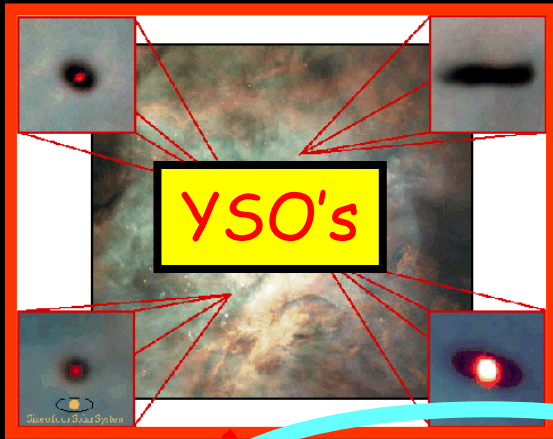
Back to cold dark space....

Refuelling the Interstellar medium.

Planet Formation

ISM

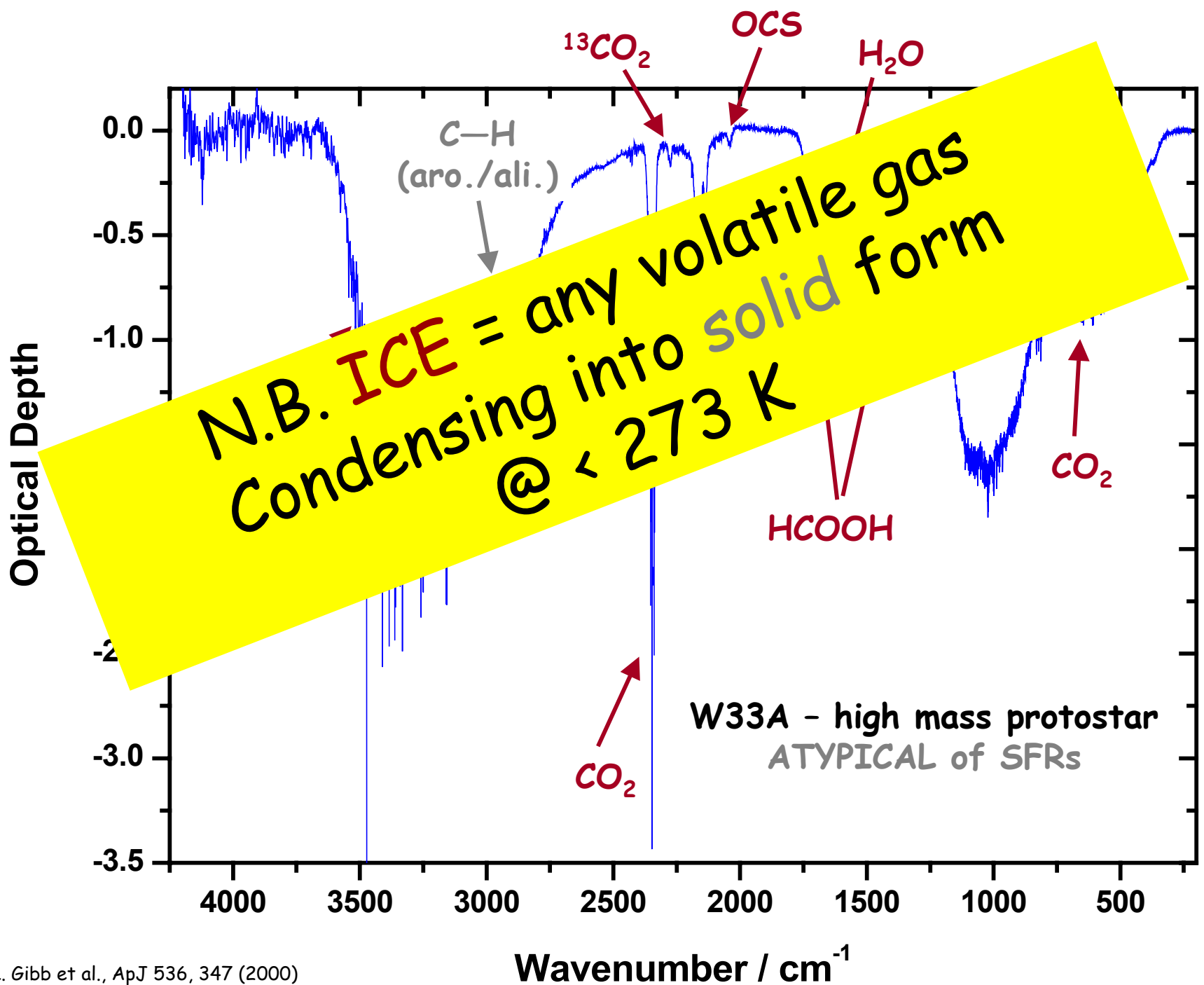
= 99% by mass GAS
(H / H₂ / He)
& 1 % DUST
(Si / C · Mg, Fe, Al, Na, K)



Formation of life

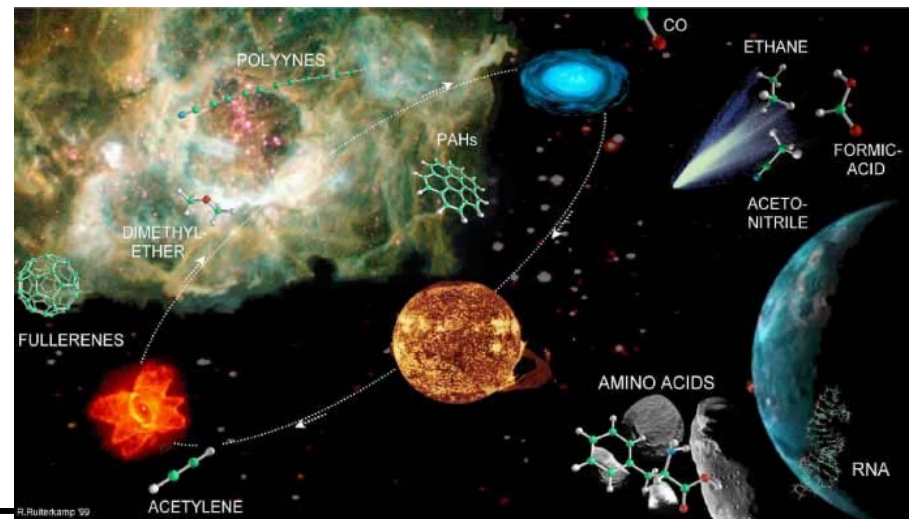
Star Death





2	3	4	5	6	7	8	9+
H ₂	C ₃	c-C ₃ H	C ₅	C ₅ H	C ₆ H	CH ₃ C ₃ N	CH ₃ C ₄ H
AlF	C ₂ H	l-C ₃ H	C ₄ H	l-H ₂ C ₄	CH ₂ CHCN	HCOOCH ₃	CH ₃ CH ₂ CN
AlCl	C ₂ O	C ₃ N	C ₄ Si	C ₂ H ₄	CH ₃ C ₂ H	CH ₃ COOH	(CH ₃) ₂ O
C ₂	C ₂ S	C ₃ O	l-C ₃ H ₂	CH ₃ CN	HC ₅ N	C ₇ H	CH ₃ CH ₂ OH
CH	CH ₂	C ₃ S	c-C ₃ H ₂	CH ₃ NC	HCOCH ₃	CH ₂ OHCHO	HC ₇ N C ₈ H
CH ⁺	HCN	C ₂ H ₂	CH ₂ CN	CH ₃ OH	NH ₂ CH ₃		CH ₃ C ₅ N
CN	HCO	CH ₂ D ⁺	CH ₄	CH ₃ SH	c-C ₂ H ₄ O		(CH ₃) ₂ CO
CO	HCO ⁺	HCCN	HC ₃ N	HC ₃ NH ⁺	CH ₂ CHOH		NH ₂ CH ₂ COO
CO ⁺	HCS ⁺	HCNH ⁺	HC ₂ NC	HC ₂ CHO			H? HC ₉ N
CP	HOC ⁺	HNCO	HCOOH	NH ₂ CHO			HC ₁₁ N
CSi	H ₂ O	HNCS	H ₂ CHN	C ₅ N			
HCl	H ₂ S	HOCO ⁺	H ₂ C ₂ O				
KCl	HNC	H ₂ CO	H ₂ NCN				
NH	HNO	H ₂ CN	HNC ₃				
NO	MgCN	H ₂ CS	SiH ₄				
NS	MgNC	H ₃ O ⁺	H ₂ COH ⁺				
NaCl	N ₂ H ⁺	NH ₃					
OH	N ₂ O	SiC ₃					
PN	NaCN						
SO	OCS						
SO ⁺	SO ₂						
SiN	c-SiC ₂						
SiO	CO ₂						

List of Detected Cosmic Molecules in Interstellar and Circumstellar Environments.



Solid-State Chemistry of Star-Forming Regions

Grain provides a 'catalytic' surface

Chemistry @ "ice" surface
e.g. CO_2 CH_3OH formation

LABORATORY STUDIES

@ surface

Disk

reaction products

H_2O

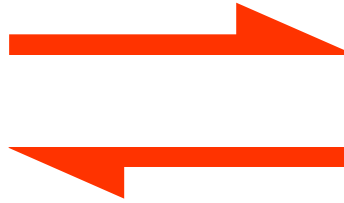
reaction product

(ects)

To understand chemistry we ALSO need to comprehend PHYSICAL behaviour of ices and surface functional groups

Chemistry in "ice" film
e.g. OCN -formation

Solid-State Astrochemistry



Astronomy

small grains

$P < 10^{-10} - 10^{-15}$ mbar
(dominated by H₂ then CO)

$T_{\text{grain}} = 10 - 300$ K

$T_{\text{gas}} = 10 - 1000$ K

1 Lyman α / Lyman-Werner band UV photon
per 10⁶ years per grain

1 atom / molecule - grain collision per 10⁴ years

1 X-Ray / CR 'direct hit' per 10⁵ years

Surface Science

To date = flat surfaces

$P < 10^{-10}$ mbar
(also dominated by H₂ then CO)

$T_{\text{grain}} = 10 - 450$ K

$T_{\text{gas}} = 100 - 300$ K

1 Lyman α / Lyman-Werner band UV photon
per molecule per second!! (≈ 5 sec \cong ISM)

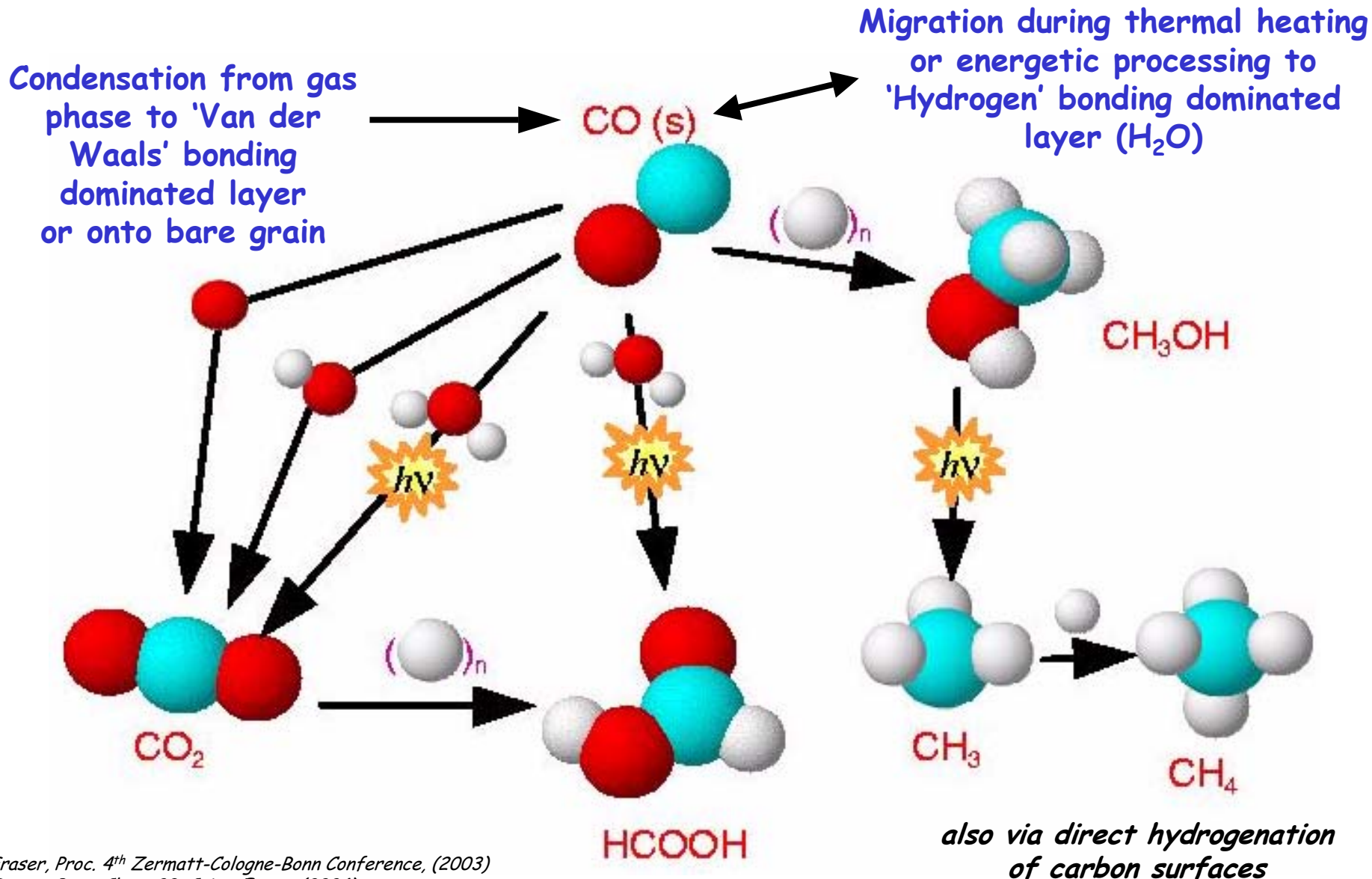
@ 1 L (Langmuir) dose = 10¹⁵ molec cm⁻² s⁻¹

1 X-Ray / CR ... ✗

S
gas
D

CR

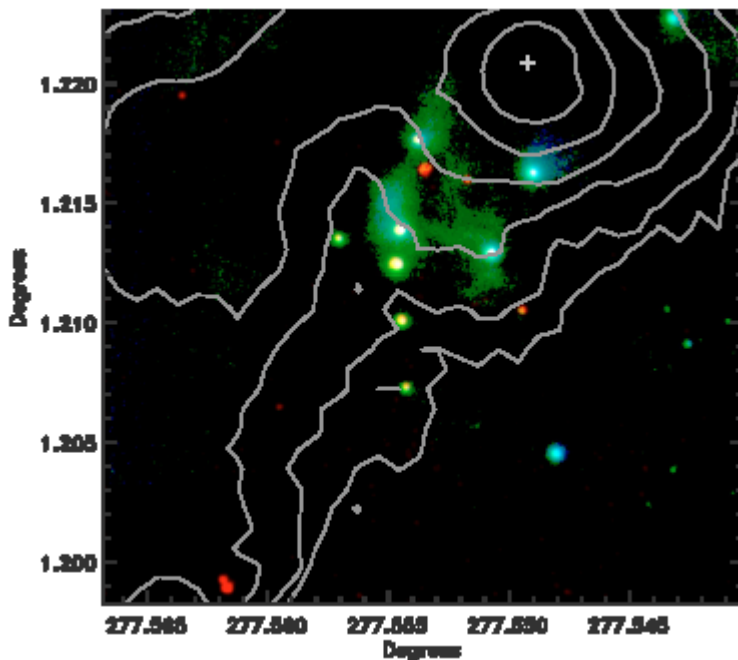
Postulated Chemistry of Solid CO in SFRs



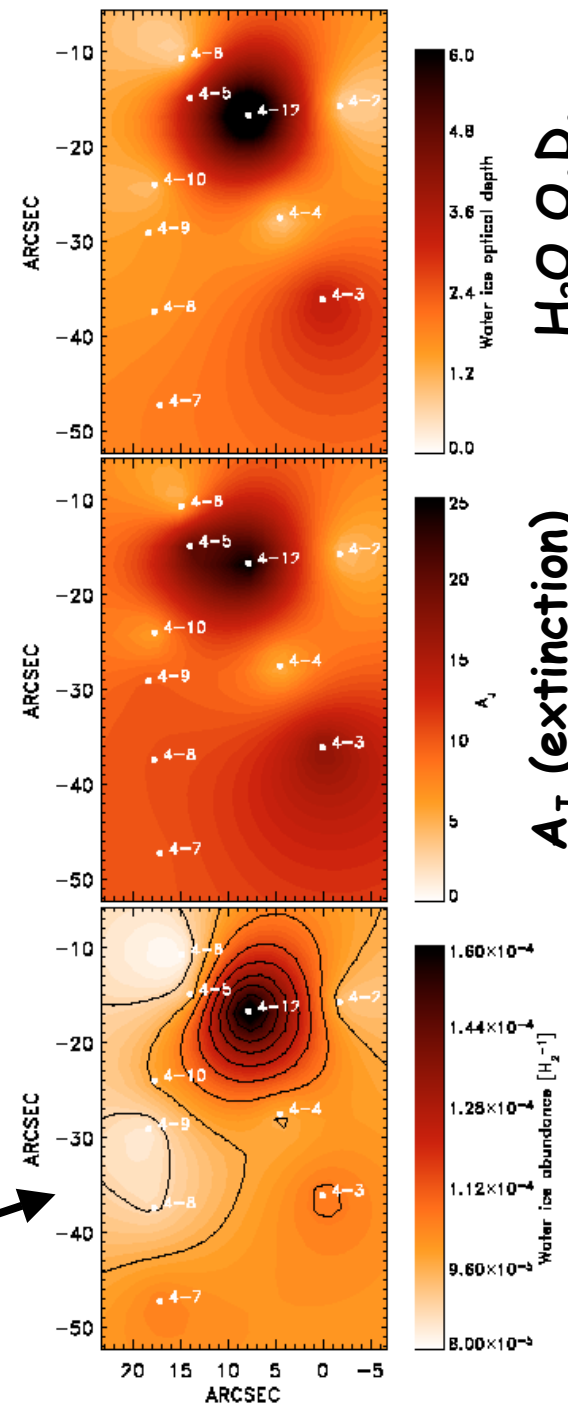
Mapping Ices Towards SVS 4 - H₂O ice mapping

N.B. proposed observational programmes on ESO VLT / SPITZER to continue mapping especially H₂O / and CO (H₂O-rich and pure) and CO₂ on SAME lines of sight

such work equally conducted in collab. CfA (Harvard) / Caltech (Keck) / JPL (Spitzer)

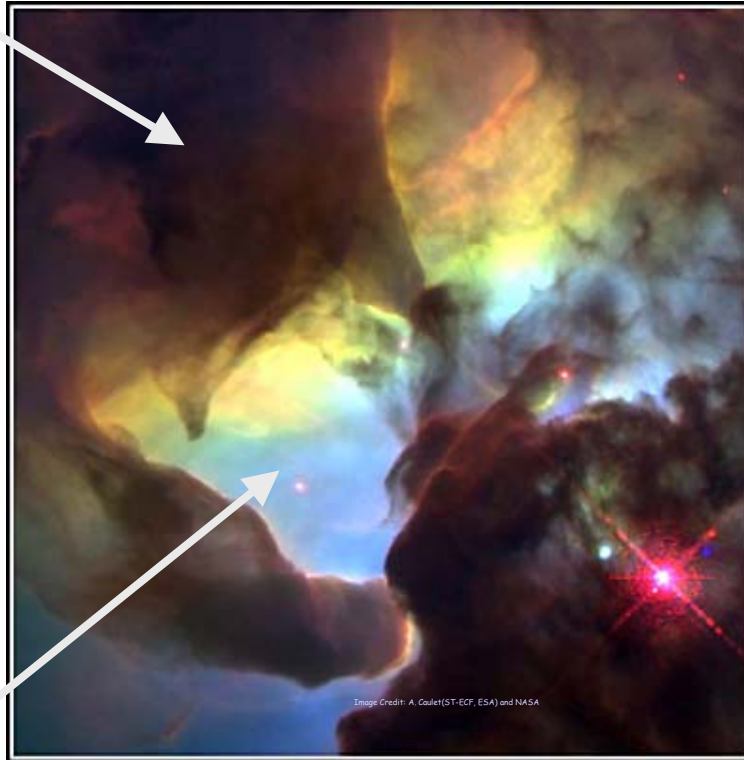


vs.. [H₂]
i.e. measure of formation efficiency



H₂O Formation - What precisely am I going to do?

H₂O (s) = most abundant 'ice'
porous & amorphous

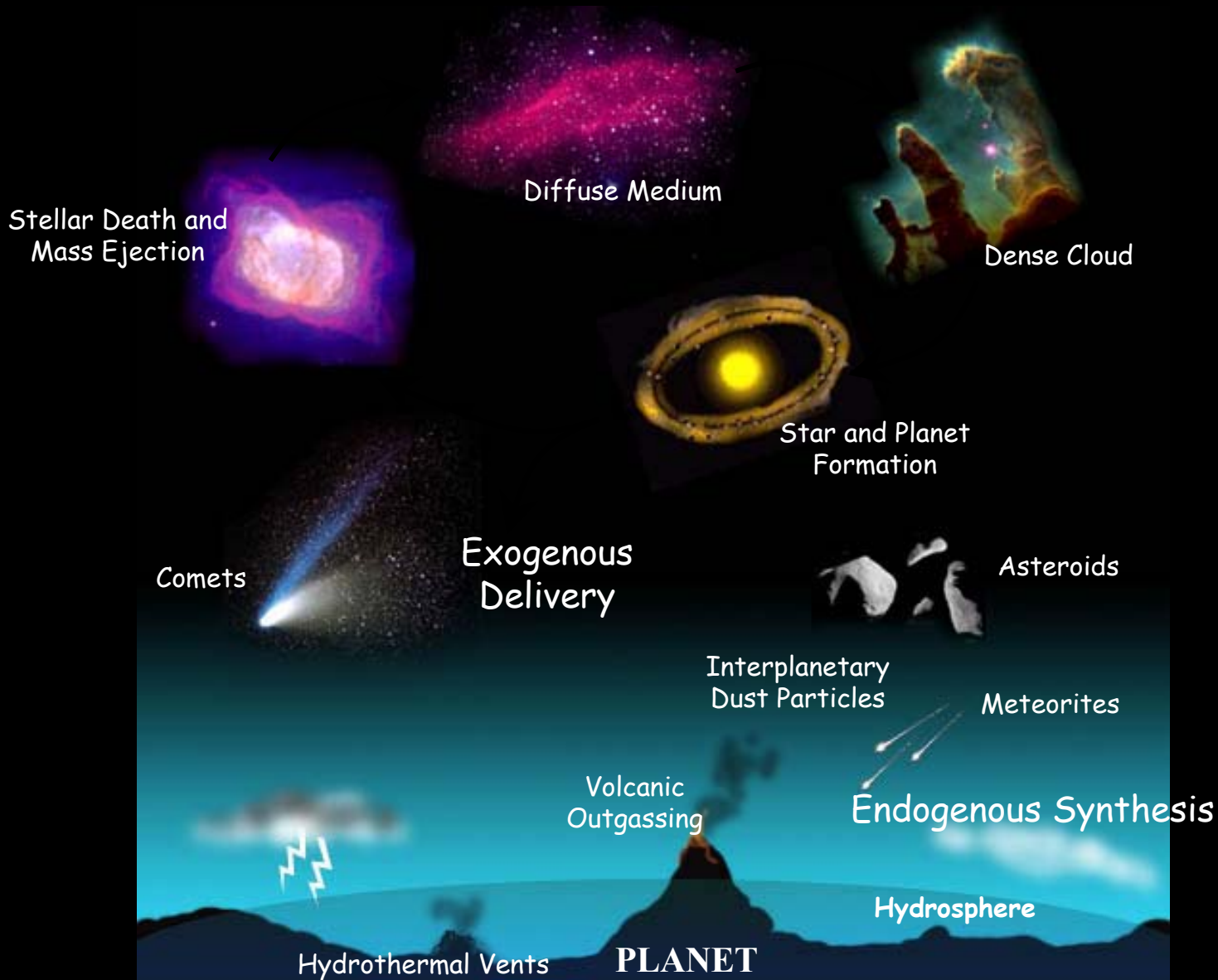


NO H₂O (g)

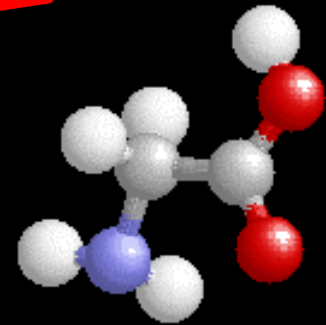
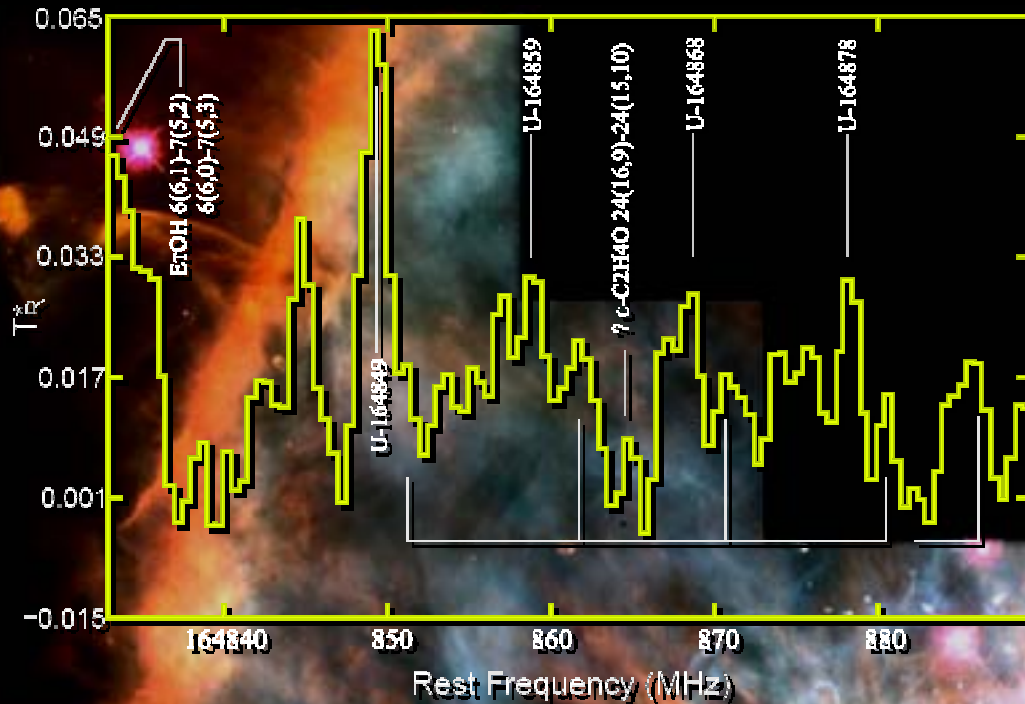
NO H₂O (s)

∴ H₂O formed on grains

So, do we make the building blocks of life in space... and then transport them to planets?



Complex molecules such as amino acids (the building blocks of life), have also been found in space

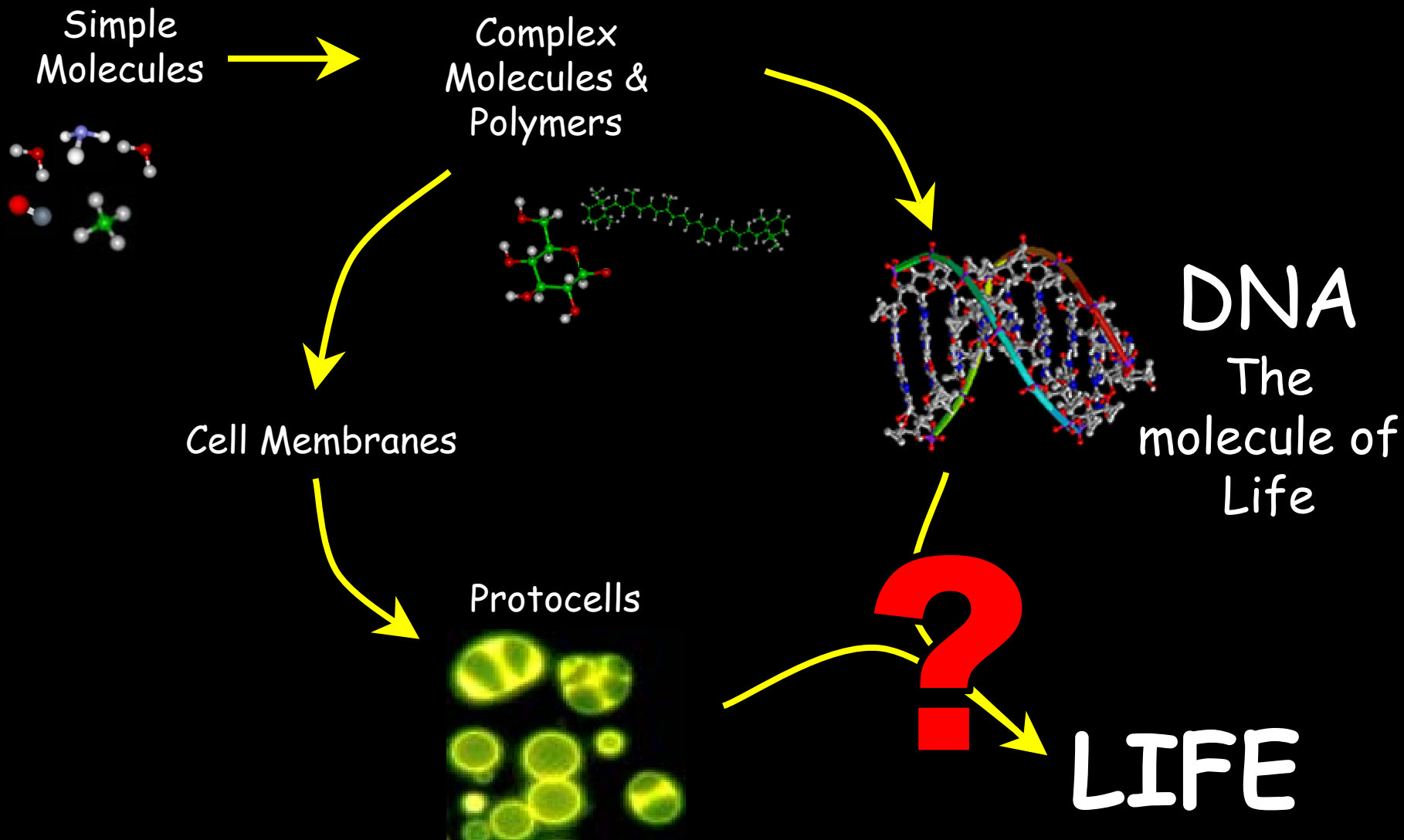


glycine
in Orion!

Credit: Y.-J. Kuan, S.B. Charnley, H.-C. Huang,
W.-L. Tseng, Z. Kiesel, *Astrophys J.* **593**, 848 (2003)

Credit: C.R. O'Dell/Rice University, NASA.

These complex molecules are important for forming the molecules of Life...



Acknowledgements



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