

# ***Lecture 8***

## ***Observational Cosmology***

### ***Parameters of Our Universe***

#### ***The Concordance Model***

# Time and Distance vs Redshift

$$\frac{d}{dt} \left( x = 1 + z = \frac{R_0}{R} \right) \rightarrow dt = \frac{-dx}{x H(x)}$$

$$\text{Friedmann: } H(x) = H_0 \sqrt{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0) x^2}$$

$$\Omega_0 \equiv \Omega_M + \Omega_\Lambda \quad (\Omega_R = 0)$$

look - back time :

$$t(z) = \int_{t_e}^{t_0} dt = \int_1^{1+z} \frac{dx}{x H(x)} = \frac{1}{H_0} \int_1^{1+z} \frac{dx}{x \sqrt{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0) x^2}}$$

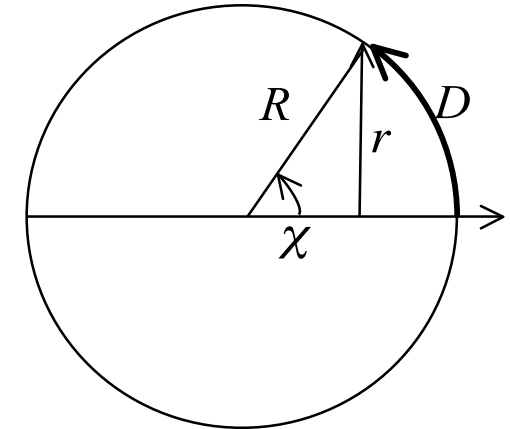
radial distance :

$$D_0(z) = R_0 \chi = \int_{t_e}^{t_0} \frac{R_0}{R(t)} c dt = \int_1^{1+z} \frac{x c dx}{x H(x)} = \frac{c}{H_0} \int_1^{1+z} \frac{dx}{\sqrt{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0) x^2}}$$

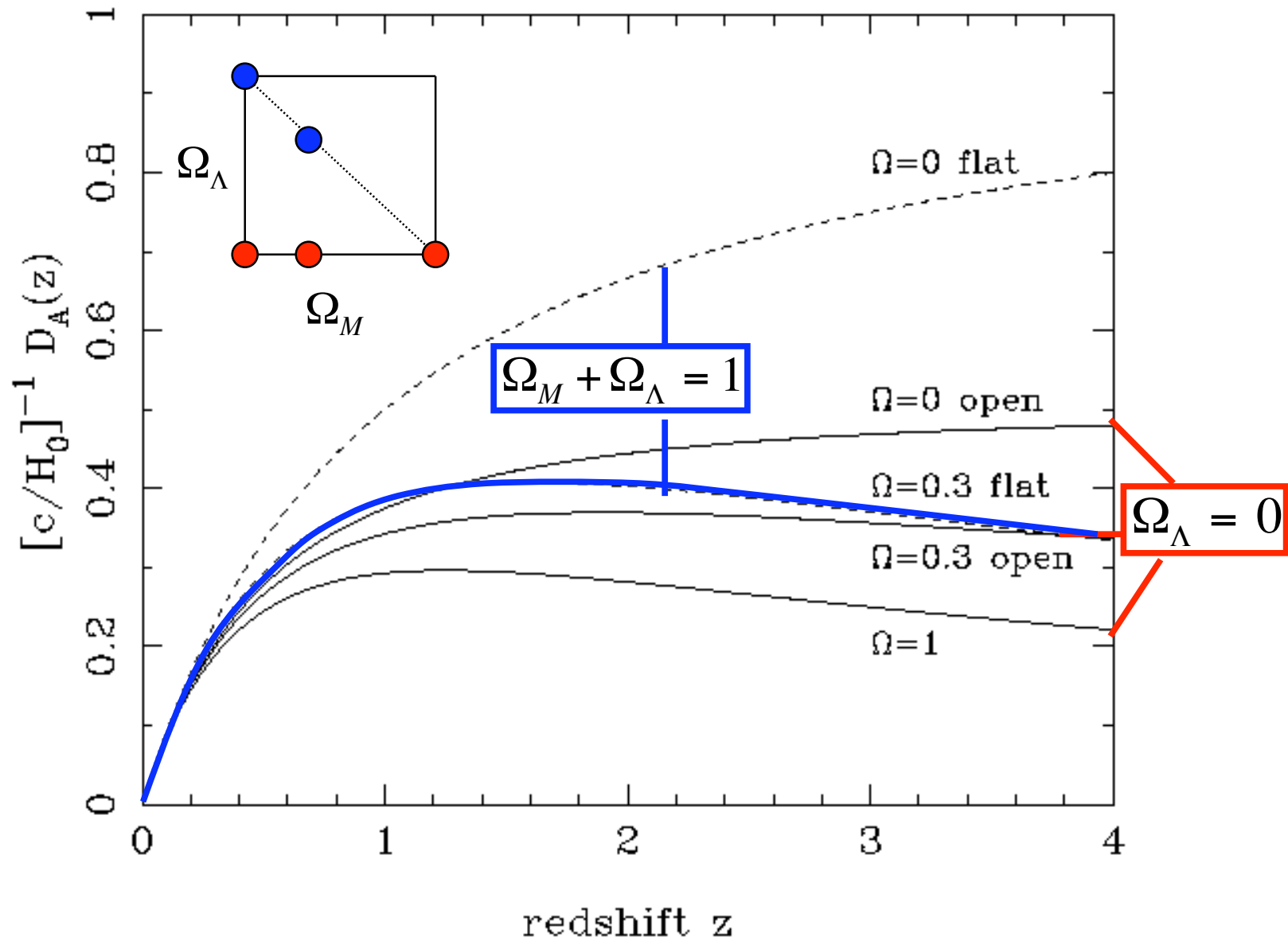
$$\text{circumferencial distance: } r_0 = R_0 S_k(\chi) \quad R_0 = \frac{c}{H_0} \left( \frac{k}{\Omega_0 - 1} \right)^{1/2}$$

$$\text{angular diameter distance: } D_A = r_0 / (1 + z)$$

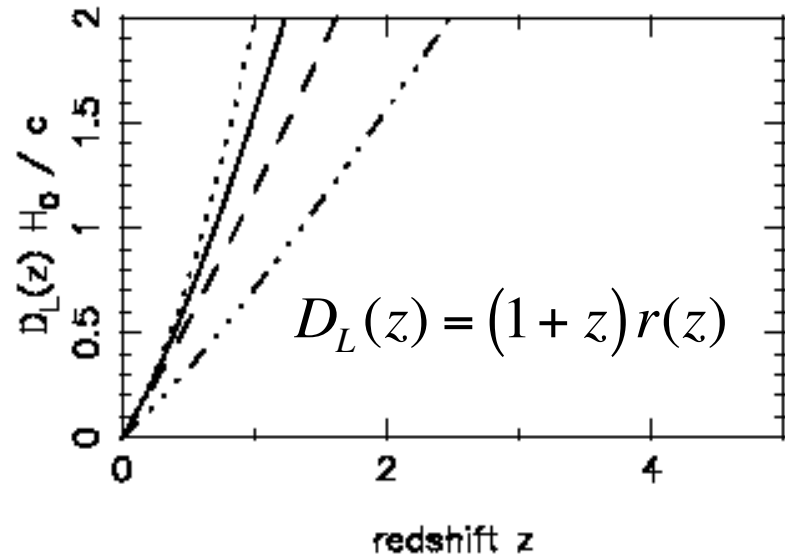
$$\text{luminosity distance: } D_L = (1 + z) r_0$$



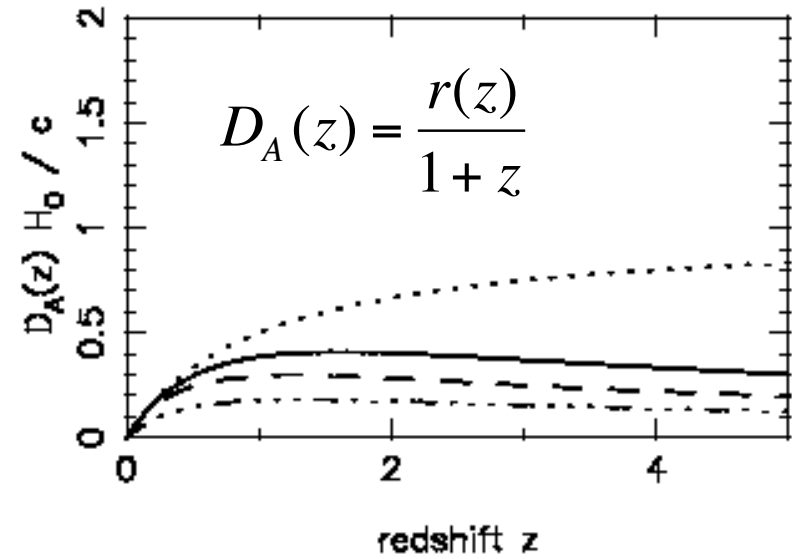
# Angular Diameter Distance



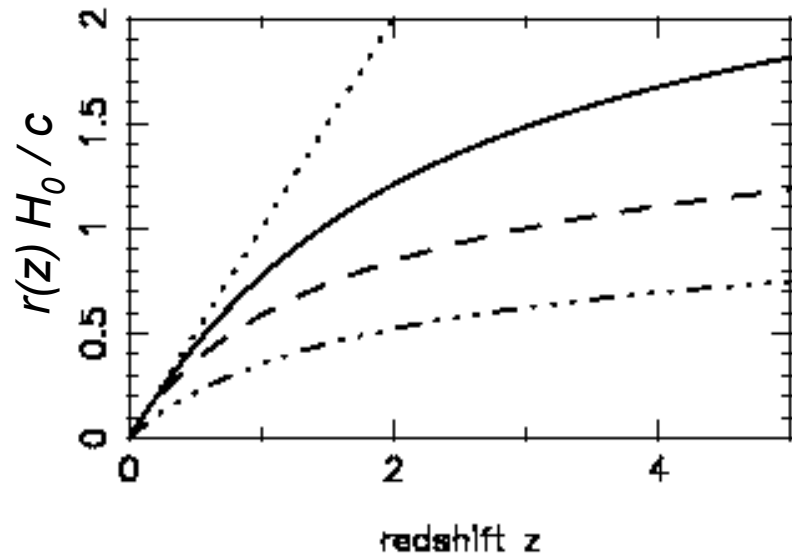
Luminosity Distance



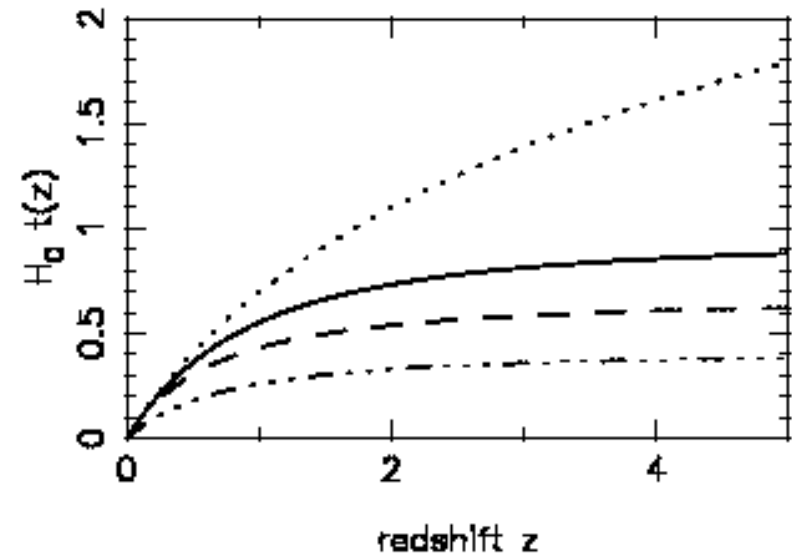
Angular Diameter Distance



$$r(z) = R_0 S_k(\chi(z))$$



Lookback Time



# “Concordance Model” Parameters

$$H_0 \equiv 100 h \frac{\text{km/s}}{\text{Mpc}} \approx 70 \frac{\text{km/s}}{\text{Mpc}} \quad h \approx 0.7$$

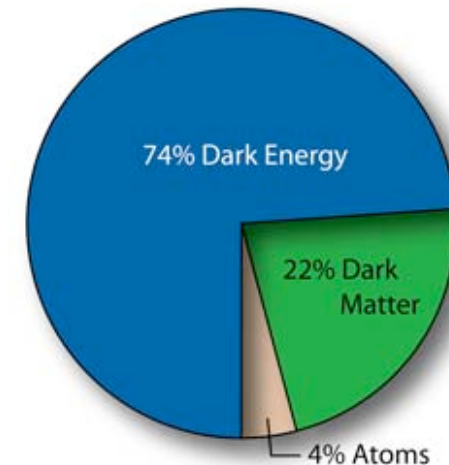
$$\Omega_R \approx 4.2 \times 10^{-5} h^{-2} \approx 8.4 \times 10^{-5} \quad (\text{CMB photons} + \text{neutrinos})$$

$$\Omega_B \sim 0.02 h^{-2} \sim 0.04 \quad (\text{baryons})$$

$$\Omega_M \sim 0.3 \quad (\text{Dark Matter})$$

$$\Omega_\Lambda \sim 0.7 \quad (\text{Dark Energy})$$

$$\Omega_0 \equiv \Omega_R + \Omega_M + \Omega_\Lambda = 1.0 \quad \rightarrow \quad \textit{Flat Geometry}$$



# Our (Crazy?) Universe

$$H_0 \approx 70 \frac{\text{km/s}}{\text{Mpc}}$$

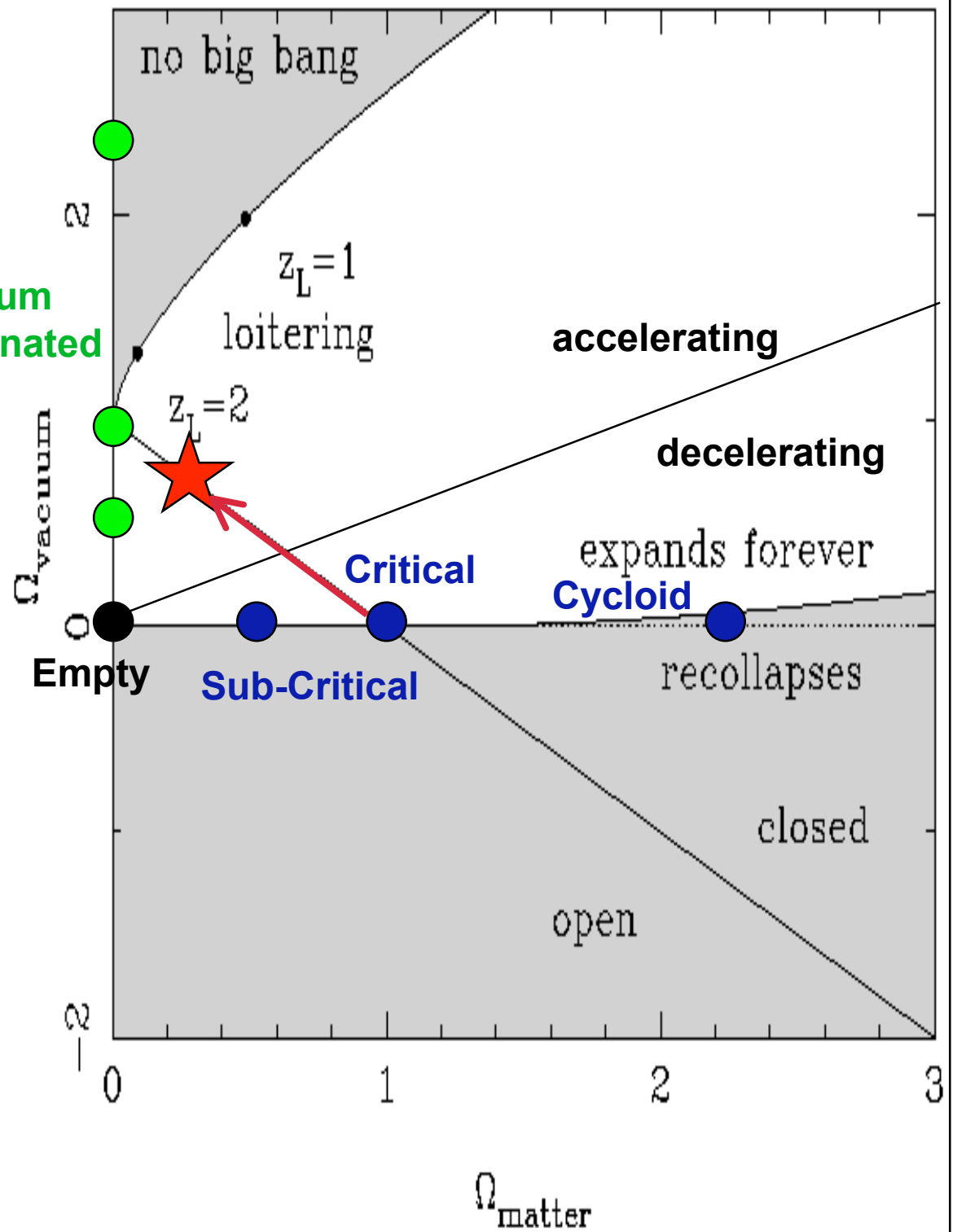
$$\Omega_M \sim 0.3$$

$$\Omega_\Lambda \sim 0.7$$

$$\Omega_R \sim 8 \times 10^{-5}$$

$$\Omega_0 = 1.0$$

Vacuum  
Dominated



# Evolution of $\Omega$

Density at a past/future epoch in units of  $\rho_c = 3 H_0^2 / 8\pi G$

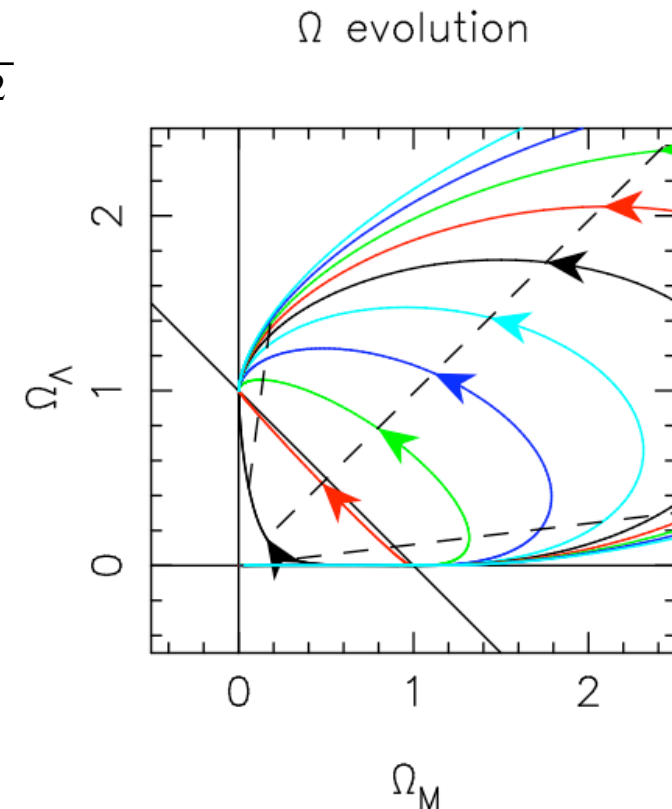
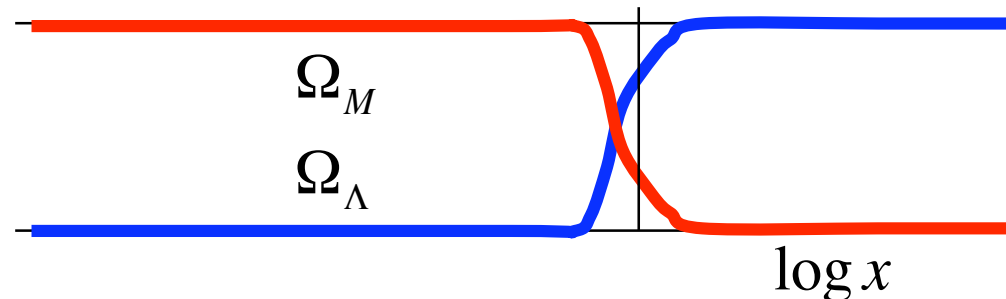
$$\Omega \equiv \frac{\rho}{\rho_c} = \sum_w \Omega_w x^{3(1+w)} = \Omega_R x^4 + \Omega_M x^3 + \Omega_\Lambda \quad x \equiv 1+z = R_0/R$$

in units of critical density  $3 H^2 / 8\pi G$  at the past/future epoch :

$$\Omega_M(x) = \frac{H_0^2}{H^2} \Omega_M x^3 = \frac{\Omega_M x^3}{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0)x^2}$$

$$\Omega_\Lambda(x) = \frac{H_0^2}{H^2} \Omega_\Lambda = \frac{\Omega_\Lambda}{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0)x^2}$$

**Do we live at a special time ?**

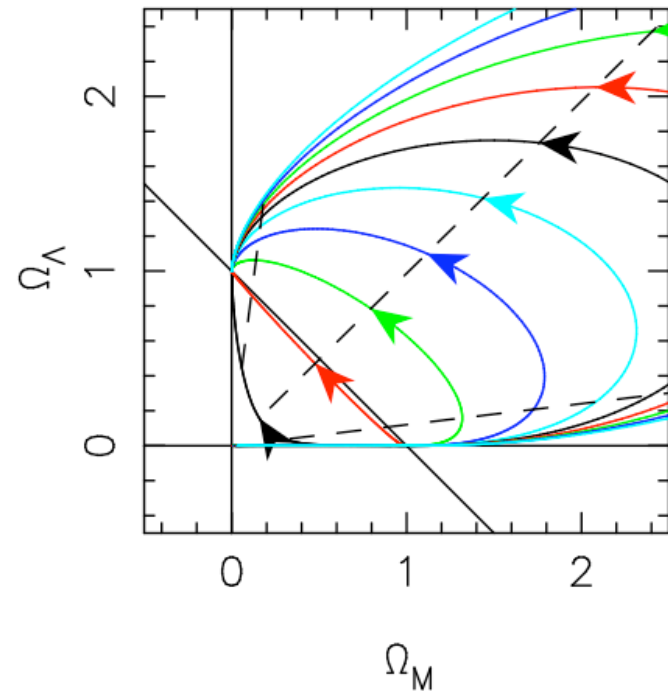
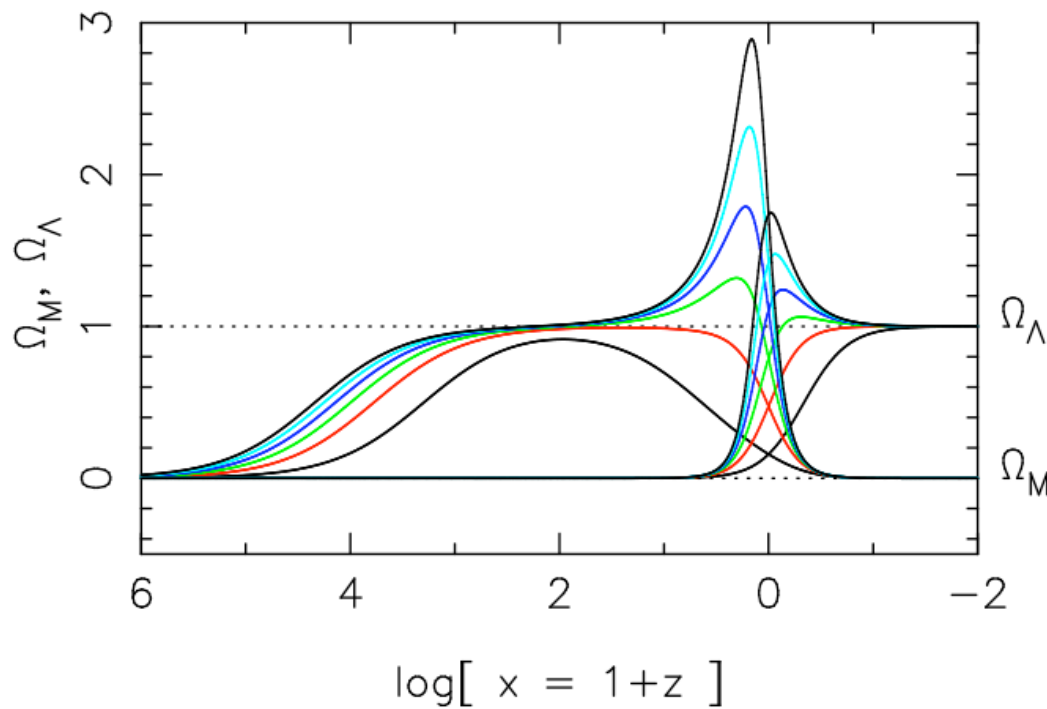


# Evolution of $\Omega$

$$\Omega_M(x) = \frac{\Omega_M x^3}{\Omega_R x^4 + \Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0)x^2}$$

$$\Omega_\Lambda(x) = \frac{\Omega_\Lambda}{\Omega_R x^4 + \Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0)x^2}$$

$\Omega$  evolution





# “Concordance” Model

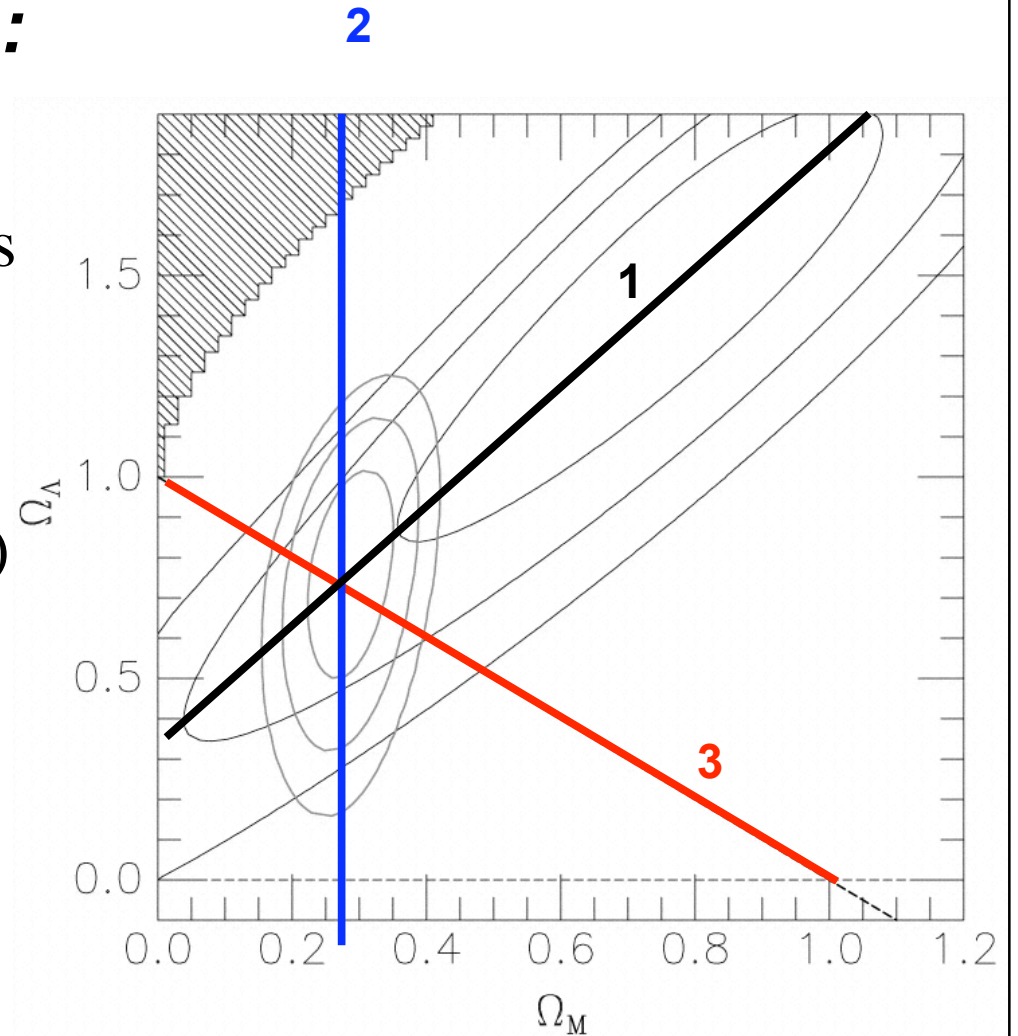
## Three main constraints:

1. Supernova Hubble Diagram
2. Galaxy Counts + M/L ratios  
 $\Omega_M \sim 0.3$
3. Flat Geometry  
(inflation, CMB fluctuations)

$$\Omega_0 = \Omega_M + \Omega_\Lambda = 1$$

concordance model

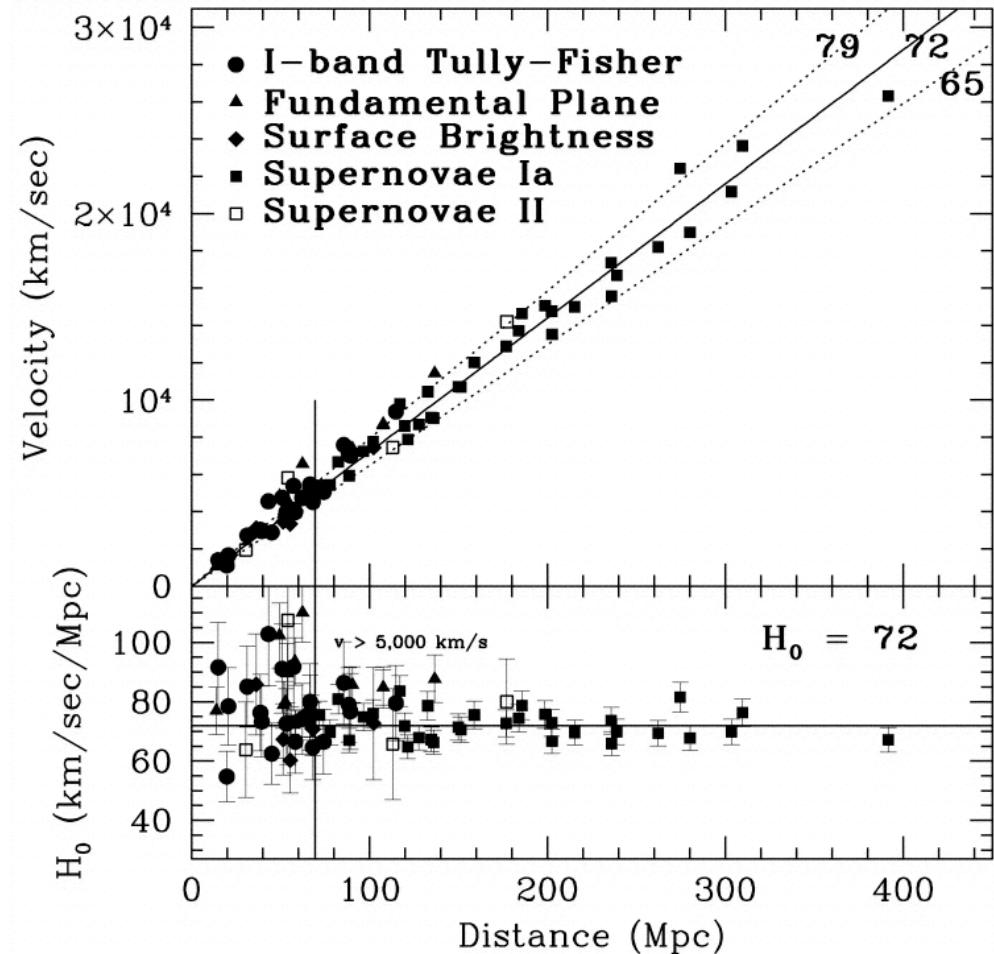
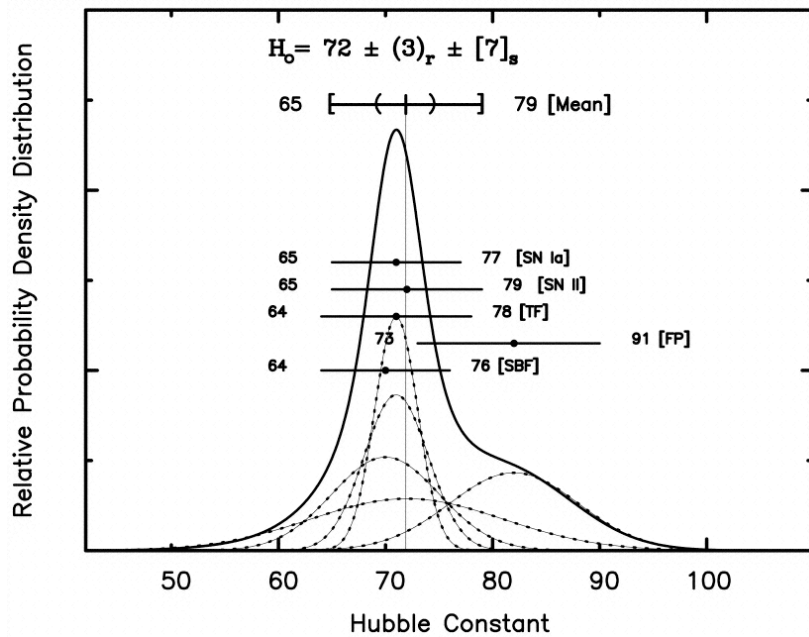
$$H_0 \approx 72 \quad \Omega_M \approx 0.3 \quad \Omega_\Lambda \approx 0.7$$



# HST Key Project

$$H_0 \approx 72 \pm 3 \pm 7 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

Freedman, et al.  
2001 ApJ 553, 47.



# *Hubble time and radius*

Hubble constant :

$$H \equiv \frac{\dot{R}}{R} \quad H_0 \equiv \left( \frac{\dot{R}}{R} \right)_0 = 100 h \frac{\text{km/s}}{\text{Mpc}} \approx 70 \frac{\text{km/s}}{\text{Mpc}}$$

Hubble time :

$$t_H \equiv \frac{1}{H_0} = 10^{10} h^{-1} \text{ yr} \approx 14 \times 10^9 \text{ yr}$$

~ age of Universe

Hubble radius :

$$R_H \equiv \frac{c}{H_0} = 3000 h^{-1} \text{ Mpc} \approx 4 \times 10^9 \text{ pc}$$

= distance light travels in a Hubble time.

~ distance to the Horizon.

# Age vs Hubble time

$$H = \frac{\dot{R}}{R} \quad H = \left. \frac{\dot{R}}{R} \right|_{t = t_0}$$

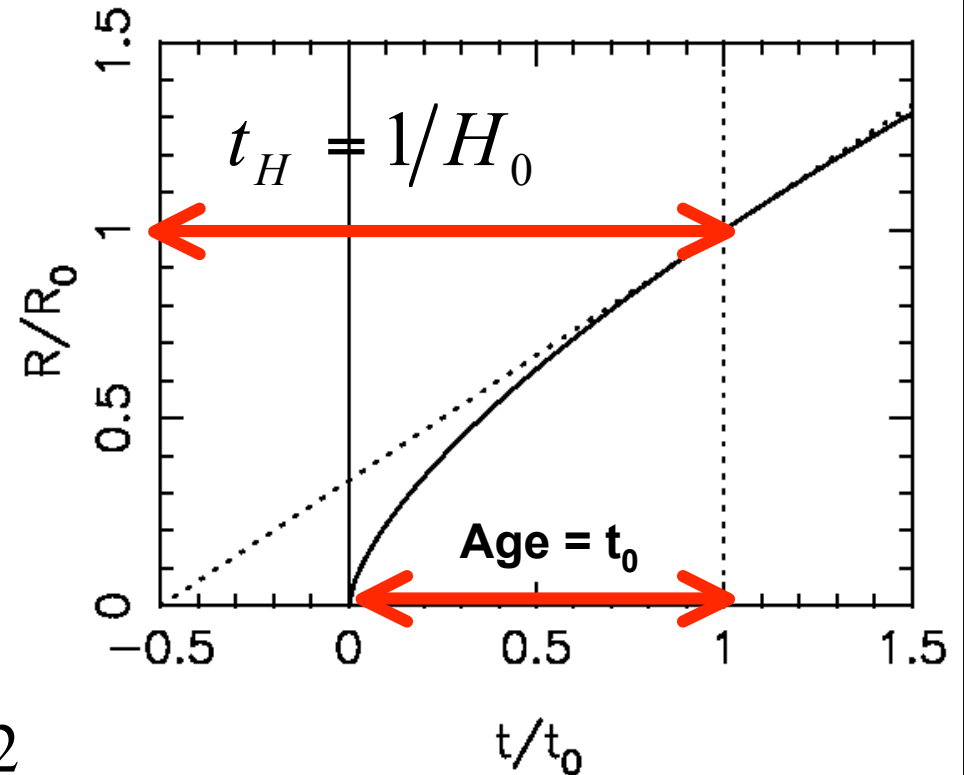
deceleration decreases age

acceleration increases age

e.g. matter dominated

$$R \propto t^{2/3} \quad \Rightarrow \quad \dot{R} = \frac{2}{3} \frac{R}{t}$$

$$H = \frac{\dot{R}}{R} = \frac{2}{3} \frac{1}{t} \quad \Rightarrow \quad t_0 = \frac{2}{3} \frac{1}{H_0} = \frac{2}{3} t_H$$



# Beyond $H_0$

- **Globular cluster ages:**  $t < t_0 \rightarrow$  acceleration
- **Radio jet lengths:**  $D_A(z) \rightarrow$  deceleration
- **Hi-Redshift Supernovae:**  $D_L(z) \rightarrow$  acceleration
  
- **Dark Matter estimates**  $\rightarrow \Omega_M \sim 0.3$
  
- **Inflation**  $\rightarrow$  Flat Geometry  $\Omega_0 \approx 1.0$
- **CMB power spectra**
  
- **“Concordance Model”**

$$\Omega_M \sim 0.3 \quad \Omega_\Lambda \sim 0.7$$

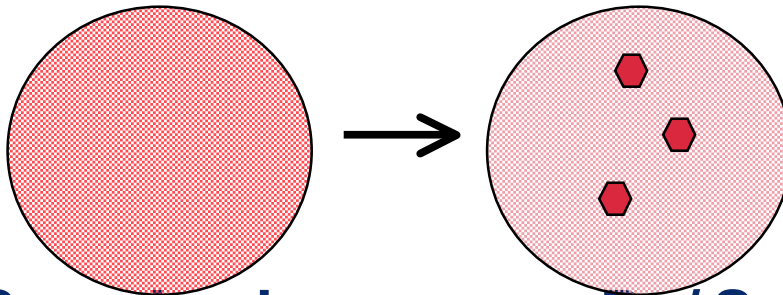
$$\Omega_0 = \Omega_M + \Omega_\Lambda \approx 1.0$$

# Age Constraints

- **Nuclear decay ( U, Th -> Pb )**
  - Decay times for ( $^{232}\text{Th}$ ,  $^{235}\text{U}$ ,  $^{238}\text{U}$ ) = (20.3, 1.02, 6.45) Gyr
  - 3.7 Gyr = oldest Earth rocks
  - 4.57 Gyr = meteorites
  - ~10 Gyr = time since supernova produced U, Th
    - (  $^{235}\text{U} / ^{238}\text{U} = 1.3 \rightarrow 0.33$ ,  $^{232}\text{Th} / ^{238}\text{U} = 1.7 \rightarrow 2.3$  )
- **Stellar evolution**
  - 13-17 Gyr = oldest globular clusters
- **White dwarf cooling**
  - ~13 Gyr = coolest white dwarfs in M4

# Nuclear Decay Chronology

- **P=parent D=daughter S=stable isotope of D**
- **Chemical fractionation changes P/S but not D/S:**



e.g. concentrate D+S in crystals

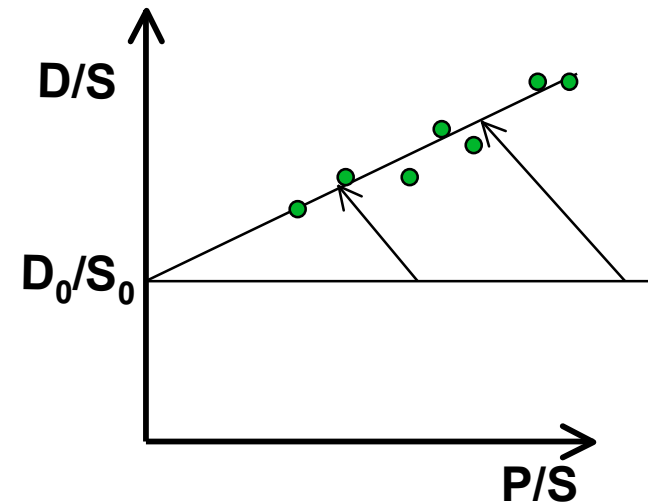
- **Samples have same  $D_0/S_0$  various  $P_0/S_0$**
- **P decays to D:**

$$P(t) = P_0 e^{-t/\tau}$$

$$D(t) = D_0 + P_0(1 - e^{-t/\tau}) \quad S(t) = S_0$$

$$= D_0 + P(t)(e^{t/\tau} - 1)$$

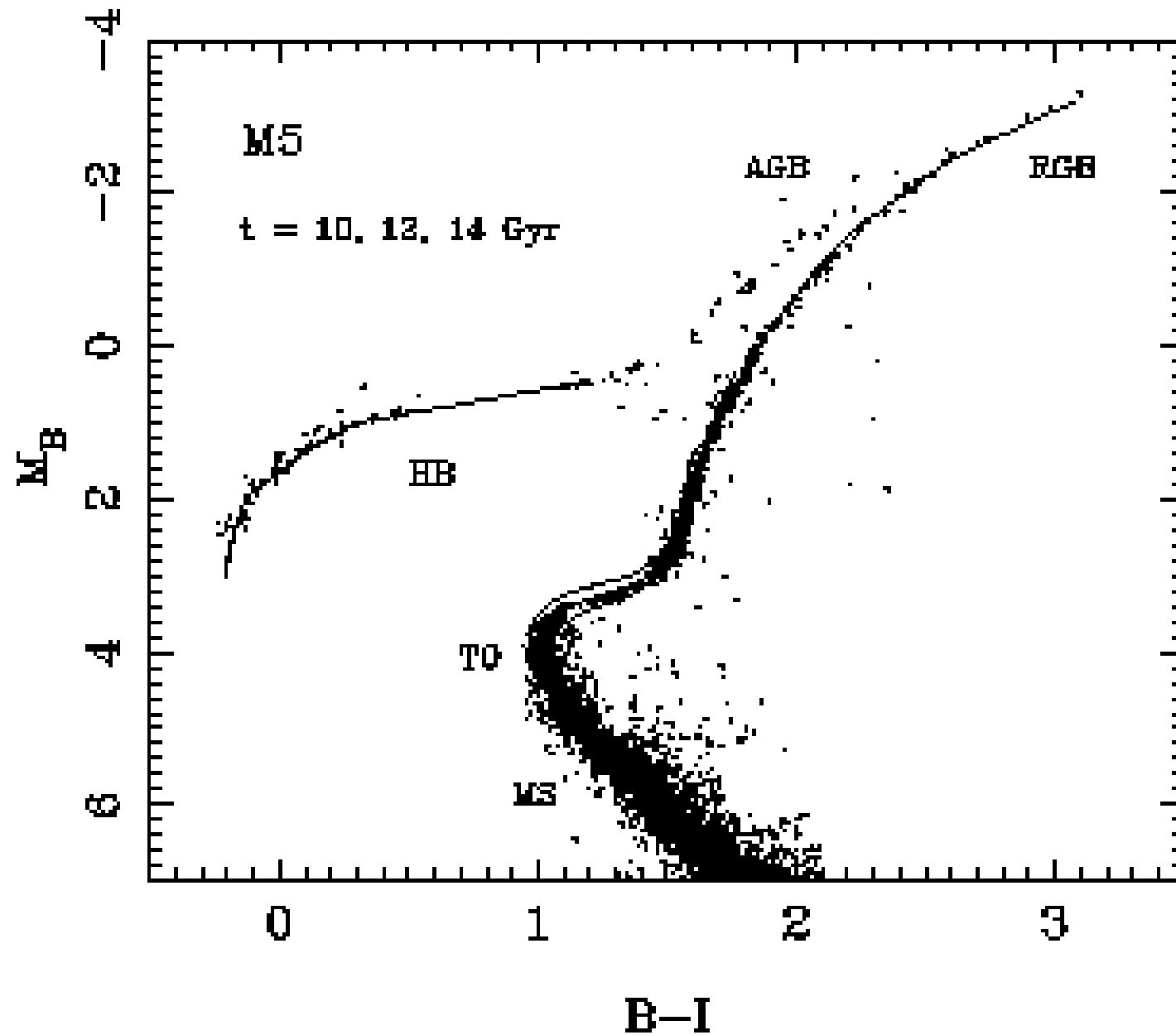
$$\frac{D(t)}{S(t)} = \frac{D_0}{S_0} + \frac{P(t)}{S_0}(e^{t/\tau} - 1)$$



Observed slope =  $(e^{t/\tau} - 1)$

gives age  $t/\tau$ , typically to  $\sim 1\%$

# *Globular Cluster Ages*

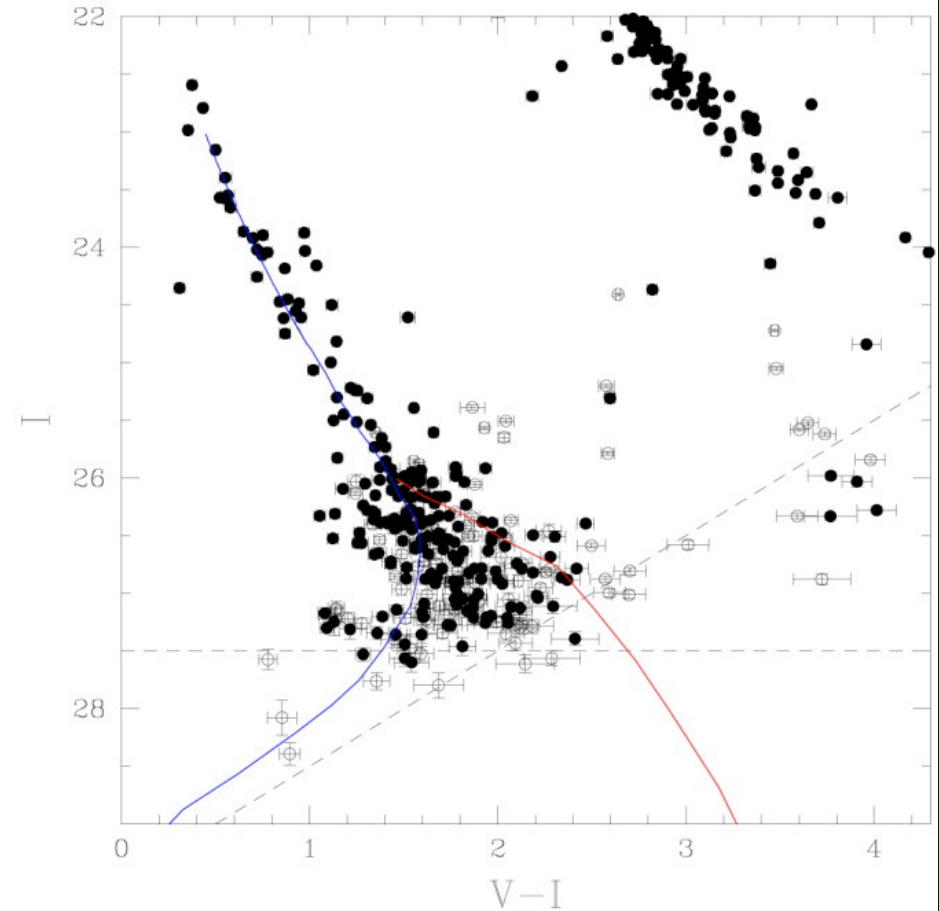
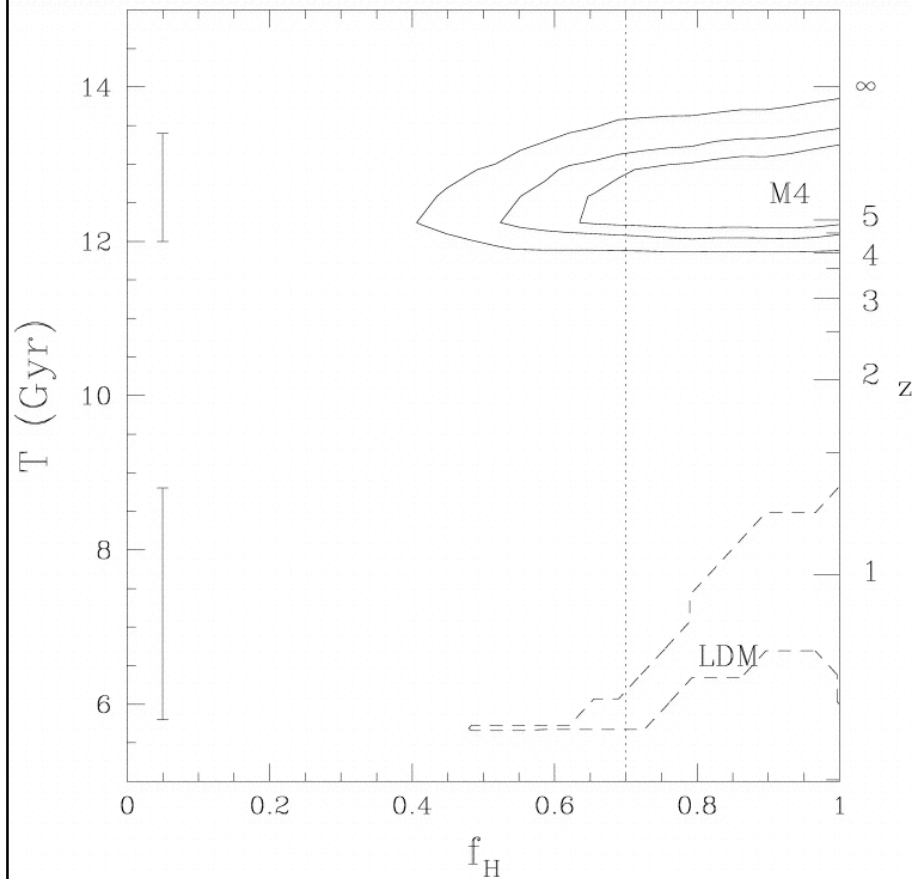




# Coollest White Dwarfs

$12.7 \pm 0.7$  Gyr

Hansen et al. 2002 ApJ 574,155



**White dwarf cooling ages  
--> star formation at  $z > 5$ .**

Cooling times have been measured using “ZZ Ceti” oscillation period changes.

# Age Crisis (~1995)

$$H_0 t_0 = \int_1^{\infty} \frac{dx}{x \sqrt{\Omega_M x^3 + \Omega_\Lambda + (1 - \Omega_0) x^2}}$$

observations :

$$H_0 = 72 \pm 8 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

$$t_0 \geq 14 \pm 2 \text{ Gyr old globular clusters}$$

$$H_0 t_0 \geq 1.0 \pm 0.15$$

**Globular clusters older than the Universe ?**

**Inconsistent with critical-density matter-only model :**

$$H_0 t_0 = \frac{2}{3} \quad \text{for } (\Omega_M, \Omega_\Lambda) = (1, 0)$$

**Strong theoretical prejudice for inflation. ( $\Omega_0 = 1$ )**  
**Doubts about stellar evolution theory (e.g. convection).**

